

## Probing freeze-in via invisible Higgs decay

Oleg Lebedev,<sup>a</sup> António P. Morais,<sup>b</sup> Vinícius Oliveira<sup>b,\*</sup> and Roman Pasechnik<sup>c</sup>

<sup>a</sup>*Department of Physics and Helsinki Institute of Physics, Gustaf Hallströmin katu 2a, FI-00014 Helsinki, Finland*

<sup>b</sup>*Departamento de Física, Universidade de Aveiro and CIDMA, Campus de Santiago, 3810-183 Aveiro, Portugal*

<sup>c</sup>*Department of Physics, Lund University, SE-223 62 Lund, Sweden*

*E-mail: [oleg.lebedev@helsinki.fi](mailto:oleg.lebedev@helsinki.fi), [aapmorais@ua.pt](mailto:aapmorais@ua.pt), [viniciuslbo@ua.pt](mailto:viniciuslbo@ua.pt), [roman.pasechnik@hep.lu.se](mailto:roman.pasechnik@hep.lu.se)*

In this work, we explore Higgs boson decays into dark matter (DM) within the context of the freeze-in mechanism under stronger coupling scenarios. Our analysis focuses on scalar DM candidates interacting through the Higgs portal, where the coupling strength can be significant without establishing thermal equilibrium, owing to a suppressed reheating temperature. We find that for DM masses as low as the MeV scale, the observed relic density can be reproduced while remaining consistent with existing collider and cosmological constraints. Notably, the framework predicts invisible Higgs decays at the level of a few percent, providing exciting prospects for experimental verification at the HL-LHC and FCC.

*2nd Training School and General Meeting of the COST Action COSMIC WISPerS (CA21106)*

*(COSMICWISPerS2024)*

*10-14 June 2024 and 3-6 September 2024*

*Ljubljana (Slovenia) and Istanbul (Turkey)*

---

\*Speaker

## 1. Introduction

The Higgs boson [1] serves as a key probe for exploring physics beyond the Standard Model (SM). Deviations in its couplings with SM particles can indicate new states, additional interactions, or exotic decay channels [2]. In this work, we investigate the Higgs decay into light DM, which manifests as an "invisible" decay mode.

In Weakly Interacting Massive Particle (WIMP) models, DM achieves thermal equilibrium with the SM plasma, and its relic density is determined by the freeze-out mechanism. For Higgs portal models [3, 4], this requires significant coupling to the Higgs field. However, LHC data constrain the branching ratio of invisible Higgs decays to below 10%, effectively excluding light thermal DM in this framework [5].

Non-thermal DM production through the freeze-in mechanism [6] offers an alternative, relying on extremely weak DM-SM couplings to avoid equilibrium. While this scenario typically evades collider constraints, the recently proposed "freeze-in at stronger coupling" [7] introduces a novel approach: if the reheating temperature  $T_R$  is lower than the DM mass  $m_{\text{DM}}$ , production and thermalization are suppressed even at larger couplings. This allows for observable Higgs decays into DM and enhances prospects for direct DM detection [8, 9].

## 2. Set-up

We assume the interactions between the SM Higgs  $\text{SU}(2)_W$ -doublet  $\mathcal{H}$  and DM are given by

$$\mathcal{L}_{hs} = \frac{\lambda_{hs}}{2} \mathcal{H}^\dagger \mathcal{H} S S, \quad (1)$$

where  $S$  is a real scalar with mass  $m_s$ . This state is assumed to be stable and thus constitute DM.

The Boltzmann equation for the DM number density  $n$  for the the *pure freeze-in* scenario takes the form

$$\dot{n} + 3Hn = 2 \Gamma(\bar{f}f \rightarrow SS), \quad (2)$$

where  $H$  is the Hubble rate, and  $\Gamma(\bar{f}f \rightarrow SS)$  is the DM production rate per unit volume. Here, the factor of two signifies production of two DM quanta in each reaction.

### 2.1 Hadronic contributions

The DM production is primarily driven by processes involving elementary fermions in the initial state. However, at temperatures below the critical QCD temperature,  $T_c \sim 150$  MeV, quarks are confined into hadrons, and the reactions instead involve hadrons and leptons. Consequently, our approach must be adapted to account for this transition.

The production rate  $\Gamma(\text{SM} \rightarrow SS)$  can be rewritten as the DM annihilation rate. Indeed, consider the reaction  $1 + 2 \rightarrow 3 + 4$ . Energy conservation implies that

$$e^{-E_1/T} e^{-E_2/T} = e^{-E_3/T} e^{-E_4/T}. \quad (3)$$

Thus, our first step is to factorize  $|\mathcal{M}(SS \rightarrow \text{SM})|^2$  using the fact that DM annihilation always proceeds via the Higgs mediator. We have

$$|\mathcal{M}_{SS \rightarrow \text{SM}}|^2 = |\mathcal{M}_{SS \rightarrow h}|^2 \frac{1}{(s - m_h^2)^2} |\mathcal{M}_{h \rightarrow \text{SM}}|^2, \quad (4)$$

where the usual spin state summation/averaging is implied, and  $m_h$  is the Higgs mass. The same relation applies to fermionic DM. Putting these ingredients together, we find

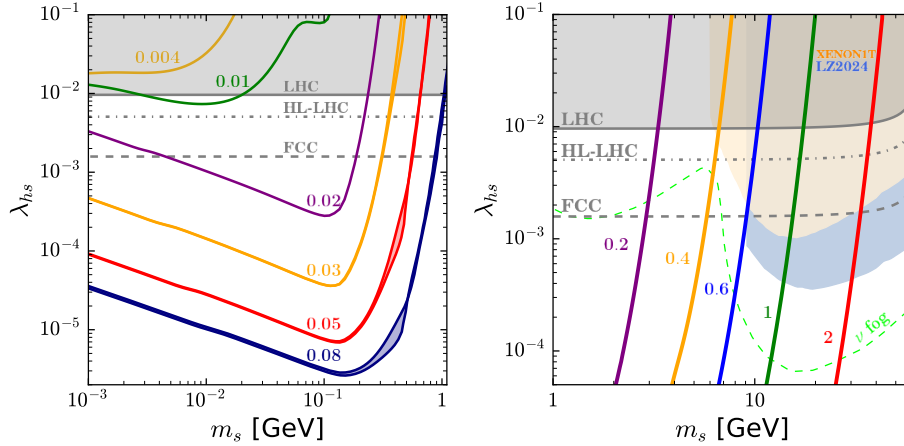
$$\Gamma_{SS \rightarrow SM}^{\text{th}} = \int \left( \prod_i \frac{d^3 \mathbf{p}_i}{(2\pi)^3 2E_i} f(p_i) \right) |\mathcal{M}_{SS \rightarrow h}|^2 \frac{1}{(s - m_h^2)^2} \times (2m_h \Gamma_h) \Big|_{m_h = \sqrt{s}}, \quad (5)$$

where the superscript ‘‘th’’ implies that DM in the initial states is treated as *thermal*. The  $\Gamma_h$  represents the Higgs decay width, this quantity is known in the SM for all Higgs masses, with variable accuracy. In our numerical analysis, we use two recent estimations by Winkler [10] and Gorbunov *et al.* [11]. Here, the SM Higgs mass in the last factor is *variable* and given by the center-of-mass energy of the annihilating DM state, in analogy with the result of [17]. We therefore conclude that at energies below the physical Higgs mass,

$$\Gamma_{SM \rightarrow SS} = \Gamma_{SS \rightarrow SM}^{\text{th}} = \frac{T}{2^5 \pi^4 m_h^4} \int_{4m_s^2}^{\infty} ds \sqrt{s(s - 4m_s^2)} K_1(\sqrt{s}/T) \Gamma_h(m_h = \sqrt{s}) |\mathcal{M}_{SS \rightarrow h}|^2 \quad (6)$$

### 3. Results and Conclusion

Our main results are shown in Fig. 1, which displays the allowed parameter space for  $\lambda_{hs}$  and  $m_s$ . The colored curves correspond to the correct relic density for different reheating temperatures  $T_R$ . For DM masses above the muon threshold,  $\lambda_{hs}$  grows exponentially with  $m_s$ . Despite the suppressed muon abundance at  $T_R < m_\mu$ , the reaction  $\bar{\mu}\mu \rightarrow SS$  remains more efficient than the electron counterpart, thanks to the larger Yukawa coupling. This effect diminishes for  $T_R < 10$  MeV. The grey shaded region is excluded by LHC constraints on invisible Higgs decays ( $\text{BR}_{\text{inv}} \lesssim 10\%$ )



**Figure 1:** Parameter space for scalar DM freeze-in. Along the colored curves marked by  $T_R$  in GeV, the correct relic density is reproduced. The shaded areas are excluded by the LHC and direct DM detection bounds. Sensitivities of the HL-LHC and FCC are shown by the grey dashed lines, while the neutrino ‘‘fog’’ for direct DM detection [12] is represented by the green dashed line.

[14]), while dashed and dash-dotted lines indicate the sensitivities of the HL-LHC ( $\text{BR}_{\text{inv}} = 3\%$  [15]) and FCC ( $\text{BR}_{\text{inv}} = 0.3\%$  [16]). The figure shows that  $T_R \gtrsim 10$  MeV is consistent with current constraints for a range of DM masses, and the corresponding parameter space can be explored

by future experiments. The DM direct detection (DD) experiments [18, 19] impose an additional constraint on the model. The strictest constraint by the recent LZ2024 result [18].

Hadronic uncertainties have limited impact on these results. At  $T_R \gtrsim 50$  MeV (but below  $T_c$ ), pion contributions slightly broaden the relic density curves. However, these differences occur in regions unlikely to be probed at colliders, minimally affecting our predictions.

**Acknowledgements.** We would like to thank Yuval Grossman and João Paulo Pinheiro for helpful comments. O.L. is grateful to the Magnus Ehrnrooth foundation for travel support, which has facilitated collaboration on this project. V.O. and A.P.M. are supported by the Center for Research and Development in Mathematics and Applications (CIDMA) through the Portuguese Foundation for Science and Technology (FCT - Fundação para a Ciência e a Tecnologia), references UIDB/04106/2020 (<https://doi.org/10.54499/UIDB/04106/2020>) and UIDP/04106/2020 (<https://doi.org/10.54499/UIDP/04106/2020>). A.P.M. and V.O. are also supported by the projects with references CERN/FIS-PAR/0019/2021 (<https://doi.org/10.54499/CERN/FIS-PAR/0019/2021>), CERN/FIS-PAR/0021/2021 (<https://doi.org/10.54499/CERN/FIS-PAR/0021/2021>) and CERN/FIS-PAR/0025/2021 (<https://doi.org/10.54499/CERN/FIS-PAR/0025/2021>). V.O. is also directly funded by FCT through the doctoral program grant with the reference PRT/BD/154629/2022 (<https://doi.org/10.54499/PRT/BD/154629/2022>). V.O. also acknowledges support by the COST Action CA21106 (Cosmic WISPerS). R.P. is supported in part by the Swedish Research Council grant, contract number 2016-05996, as well as by the European Research Council (ERC) under the European Union’s Horizon 2020 research and innovation programme (grant agreement No 668679). R.P. also acknowledges support by the COST Action CA22130 (COMETA).

## References

- [1] P. W. Higgs, Phys. Rev. Lett. **13**, 508-509 (1964).
- [2] S. D. Bass, A. De Roeck and M. Kado, Nature Rev. Phys. **3**, no.9, 608-624 (2021).
- [3] V. Silveira and A. Zee, Phys. Lett. B **161**, 136-140 (1985).
- [4] B. Patt and F. Wilczek, [arXiv:hep-ph/0605188 [hep-ph]].
- [5] A. Djouadi, O. Lebedev, Y. Mambrini and J. Quevillon, Phys. Lett. B **709**, 65-69 (2012).
- [6] L. J. Hall, K. Jedamzik, J. March-Russell and S. M. West, JHEP **03**, 080 (2010).
- [7] C. Cosme, F. Costa and O. Lebedev, Phys. Rev. D **109**, no.7, 075038 (2024).
- [8] N. Koivunen, O. Lebedev and M. Raidal, [arXiv:2403.15533 [hep-ph]].
- [9] G. Arcadi, F. Costa, A. Goudelis and O. Lebedev, JHEP **07**, 044 (2024).
- [10] M. W. Winkler, Phys. Rev. D **99**, no.1, 015018 (2019).
- [11] D. Gorbunov, E. Kriukova and O. Teryaev, [arXiv:2303.12847 [hep-ph]].

- [12] J. Billard, M. Boulay, S. Cebrián, L. Covi, G. Fiorillo, A. Green, J. Kopp, B. Majorovits, K. Palladino and F. Petricca, *et al.* Rept. Prog. Phys. **85**, no.5, 056201 (2022).
- [13] P. Gondolo and G. Gelmini, Nucl. Phys. B **360**, 145-179 (1991).
- [14] G. Aad *et al.* [ATLAS], Phys. Lett. B **842**, 137963 (2023).
- [15] P. A. Rivadeneira Bracho, “Search for invisible decays of the Higgs boson produced via vector boson fusion at the ATLAS detector with 139 fb<sup>-1</sup> of integrated luminosity,” *PhD thesis, University of Hamburg, 2022.*
- [16] P. Giacomelli, *talk at ICHEP 2018*, [https://indico.cern.ch/event/686555/contributions/2971566/attachments/1682031/2703684/Higgs-measurements-FCC-ICHEP-2018\\_169.pdf](https://indico.cern.ch/event/686555/contributions/2971566/attachments/1682031/2703684/Higgs-measurements-FCC-ICHEP-2018_169.pdf)
- [17] C. P. Burgess, M. Pospelov and T. ter Veldhuis, Nucl. Phys. B **619**, 709-728 (2001).
- [18] J. Aalbers *et al.* [LZ Collaboration], [arXiv:2410.17036 [hep-ex]].
- [19] E. Aprile *et al.* [XENON], Phys. Rev. Lett. **121**, no.11, 111302 (2018).