Study of hadronic cross-sections with cosmic ray Carbon from GeV to TeV by the DAMPE experiment

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The DArk Matter Particle Explorer (DAMPE) is a satellite-borne particle detector launched on December 17th, 2015, with different scientific objectives, looking for signatures of Dark Matter decay or annihilation, performing gamma-ray astronomy and providing precise measurements of galactic Cosmic Ray (CR) energy spectra. Accurate measurements of hadronic interaction cross sections play a key role in the determination of CR fluxes. The survival probabilities have been implemented to study hadronic interaction cross sections with the BGO calorimeter target for Carbon nuclei in a wide kinetic energy range from a few GeV to TeV, by using data collected by the DAMPE experiment. The results have been then compared with Geant4 simulations of the interaction cross sections are here presented and discussed.

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10 1. Introduction

The DArk Matter Particle Explorer (DAMPE)[1] satellite is a Chinese space mission designed 11 to study cosmic rays(CRs), gamma rays, and electrons in space. It was launched into a Sun-12 synchronous orbit around the Earth at an altitude of 500 km on December 17, 2015, and operated 13 until now. The primary goal of DAMPE is to indirectly search for dark matter by studying the 14 high-energy cosmic rays and gamma rays in space. It focuses on the detection and measurement 15 of electrons and photons[2] with high precision and energy resolution. The DAMPE also measures 16 CRs protons, helium and other nuclei[3-5] to understand the origin, propagation and acceleration 17 mechanisms by studying the properties of these CRs. The hadronic interaction cross sections play 18 a key role in the determination of cosmic ray(CR) fluxes. In this work we present the analysis 19 procedure for cross section on BGO target and specifically the survival probability measured by 20 means of flight data(FD) and simulation models. 21

22 **2. DAMPE detector**

The DAMPE detector [6] consists of a Plastic Scintillator strip Detector (PSD) [7] for charge measurement of incident particles, a Silicon-Tugsten tracKer (STK) [8] to reconstruct the trajectories of the particles, a Bismuth Germanium Oxide(BGO) [9] calorimeter to provides the measurement of energy and a powerful electron-hadron discrimination, and a NeUtron Detector (NUD).

²⁷ DAMPE's BGO calorimeter is a total absorption calorimeter. It is composed of 308 BGO ²⁸ crystal bars, which enable accurate and detailed three-dimensional imaging of the shape of particle ²⁹ showers. The BGO calorimeter in DAMPE comprise 14 layers, providing 32 radiation lengths and ³⁰ 1.6 nuclear interaction lengths. Each layer consists of 22 BGO crystal bars, with dimensions of ³¹ $25 \times 25 \times 600 \text{ mm}^3$. These crystal bars are arranged within the calorimeter and serve to measure the ³² energy deposited by particles. The key parameters of the BGO calorimeter are listed in Table I.

The FD used for the analysis described in this work were taken from January 1st, 2016 to December 31st, 2022. The number of collected events per day is about 5 millions, and the FD taking mode of DAMPE is very stable[10]. The Monte Carlo(MC) simulation for carbon and oxygen nuclei is based on Geant4[11] that is the main simulation engine for DAMPE experiment. The model of Geant4 used for this analysis is FTFP_BERT(Hadronic Model) and the energy range is from 10 GeV to 100 TeV. The FLUKA[12] simulation software has been also used to generate MC data from 100 GeV to 100 TeV.

Active area	$60 \times 60 cm^2(on - axis)$
Radiation lengths	$32X_0$
Nuclear interaction length	$\sim 1.6\lambda_I$
Longitudinal segmentation	14 layers

Table 1: BGO specifications

3. Measurement of survival probability

The BGO calorimeter below the PSD and STK is composed by weight of 56.47% Bismuth, 14.73% Germanium, 16.21% Oxygen, 9.2% Carbon elements and small amount of other elements . When N_{in} incoming nuclei enter the BGO, and N_{out}^{sur} nuclei survive without hadronic interaction, their ratios is the survival probability ε_{sur} , and depends on the amounts of materials and interaction cross sections in exponential form[13]. It can be estimated in the following way:

$$\varepsilon^{sur} = \frac{N_{out}^{sur}}{N_{in}} = exp(-n\sigma_{had}) \tag{1}$$

where n refers to the number of targets per area, which is the weight average of different path lengths of particle through BGO layers on incident direction. σ_{had} is the hadronic interaction cross sections on BGO target.

3.1 Measurement of carbon survival probability

The CRs pass through DAMPE from PSD to BGO. The following selection criteria are applied in order to select the samples. The first is a fiducial cut, which decides the effective geometric volume of the DAMPE detector from the top PSD Layer to the bottom BGO Layer. The trackfinding algorithm reconstructs the precise track in the STK to calculate the path length in the BGO. The charge measurement is mainly provided by the PSD as shown in Fig.1. To further reduce the contamnation of other nuclei, like boron, the STK and BGO first layer are uesd to supply the charge measurement too. Fig.2 shows the charge distributions due to BGO measurement for FD and MC. The N_{in} was given by:

$$N_{in}^{Z=6} = N_{Z=6}^{PSD} \cap N_{Z=6}^{STK} \cap N_{Z=6}^{BGO_{-1}st} layer$$
(2)

The contamination of carbon $N_{in}^{Z=6}$ from other nuclei was estimated to be ~2% for deposited 58 energies around 100 GeV and $\sim 10\%$ around 50 TeV. The N_{out}^{sur} is further selected by BGO second 59 and third layers to measure the survival probability that means no hadronic interaction if charge 60 doesn't change. Fig.3 a) shows the measured carbon survival probabilities between second and third 61 of BGO as functions of kinetic energy for FD and different simulation models, as well as their ratio. 62 There are a significant discrepancies between MC and FD, but the discrepancies between FLUKA 63 and FD is smaller than that between Geant4 and FD. Due to the insufficient absolute resolution of 64 the BGO charge, we cannot fully distinguish events without hadronic from events where the CR 65 interacts in BGO. 66

3.2 Measurement of oxygen survival probability

The survival probabilities of oxygen (Z=8) was analyzed in the same way. Fig.3 b) shows the measured survival probabilities of oxygen for FD and MC, as well as their ratio of MC to FD.

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Figure 1: Spectrum of the charge from 0 e to 10 e. A significant suppression of the proton is due to selection of STK.



Figure 2: Spectrum of 1^{st} BGO layer charge of FD was selected Z=6 by PSD and STK.



Figure 3: The survival probabilities of a) C and b) O for FD (black circles), Geant4 (green triangles) and FLUKA (blue squares). The difference between simulation models and FD expressed by the ratio as a function of kinetic energy (bottom figures).

Measurement of cross sections 4. 70

The inelastic hadronic cross sections σ_{had} of carbon and oxygen for Geant4 (FTFP_BERT) and 71 FLUKA are shown Fig.4. The inelastic cross sections of FLUKA is larger than Geant4 of ~ 1 % and 72 for oxygen around 100GeV is the opposite. The measured MC survival probabilities reflect the fact 73 that the survival probabilities is inversely correlated with the cross sections. Particularly for oxygen, 74 both models exhibit a tendencies consistent with the measured survival probabilities, with the cross 75 sections showing opposite performance at low and high energy. So, the DAMPE measured survival 76 probabilities measured can describe the physical variable of hadronic cross sections according to 77 the formula. 78

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1.68

1.54

1.02

1.015

1.0 1.005

0.995 0.99

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tatio FLUKA/Geant4

$$\zeta_{ratio} = \frac{\sigma_{BGO}^{FD}}{\sigma_{RGO}^{MC}} = \frac{ln(\varepsilon_{BGO}^{FD})}{ln(\varepsilon_{RGO}^{MC})}$$
(3)

FLUKA

Geant4

Fig.5 shows the FD/MC ratios of hadronic cross sections of carbon and oxygen with systematic 80 uncertainties. The cross sections of FD are even higher than the difference between the simulation 81 models. The main systematic uncertainty is from method of this study and evaluated using MC. 82 83

1.88

8.1.u 1.84 1.82 1.82

1.76 1.74 1.72 1.72 1.7

1.68

1.02

1.0

0.99 0.98

0.97

10²

10³ Kinetic er

(b)

nerav (GeV)

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Ratio FLUKA/Geant4

1.8 ະ ເງິ 1.78

The inelastic hadronic cross sections of proton and helium presented in Ref. [14].

FLUKA

Geant4

10³ Kinetic energy (GeV)

(a)

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Figure 5: The ratio of inelastic hadronic cross sections on BGO target between FD and simulation models for a) C, b) O as functions of kinetic energy for simulation models (I shift X-axis of FLUKA to make it easier to distinguish).

5. Summary

In this work, we measured the survival probabilities of carbon and oxygen in BGO calorimeter and they results to be greatly consistent with the hadronic cross sections. The cross sections of FD is larger than simulation models with significant systematic uncertainties.

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100 References

- [1] J.Chang et al., (DAMPE Collaboration) *The DArk Matter Particle Explorer mission*. Astroparticle Physics 95 (2017) 6-24.
- [2] G.Ambrosi et al., (DAMPE Collaboration) *Direct detection of a break in the teraelectronvolt cosmic-ray spectrum of electrons and positrons.* Nature, 552.7683 (2017): 63-66
- [3] Q.An et al., (DAMPE Collaboration) *Measurement of the cosmic ray proton spectrum from 40 GeV to 100 TeV with the DAMPE satellite.* Science advances, 5.9 (2019): eaax3793.
- [4] F. Alemanno et al., (DAMPE Collaboration) *Measurement of the cosmic ray helium energy spectrum from 70 GeV to 80 TeV with the DAMPE space mission*. APhysical Review Letters,
 126.20 (2021): 201102.
- [5] F. Alemanno et al., (DAMPE Collaboration)*Detection of spectral hardenings in cosmic-ray boron-to-carbon and boron-to-oxygen flux ratios with DAMPE.* Science Bulletin, 67.21
 (2022): 2162-2166
- [6] Ambrosi G et al., *The on-orbit calibration of DArk Matter Particle Explorer*. Astroparticle
 Physics, 106 (2019): 18-34.
- [7] Y. Yu et al., *The Plastic Scintillator Detector For Dampe*. Astroparticle Physics, 94 (2017):
 1-10.
- [8] P. Azzarello et al., *The DAMPE silicon–tungsten tracker*. Nucl. Instrum. Meth. A 831 (2016): 378-384
- [9] Wu L et al., *Calibration and Status of the 3-D Imaging Calorimeter of DAMPE for Cosmic Ray Physics on Orbit.* IEEE Transactions on Nuclear Science, 65.8 (2018) : 2007-2012.
- [10] C. Liu et al., *Study on the aging of the BGO calorimeter of the DAMPE experiment in space*.
 IEEE Transactions on Nuclear Science, doi: 10.1109/TNS.2023.3262399.
- [11] Agostinelli C et al, *GEANT4-a simulation toolkit*. Nucl.Instrum.Meth. A 506.3 (2003):250 303
- [12] T.T. Böhlen et al., *The FLUKA Code: Developments and Challenges for High Energy and Medical Applications.* Nuclear Data Sheets, 120 (2014):211-214
- ¹²⁷ [13] Q. Yan et al., *Measurements of nuclear interaction cross sections with the Alpha Magnetic* ¹²⁸ Spectrometer on the International Space Station. Nuclear Physics A, 996 (2020):121712
- [14] P. Coppin et al., *Probing hadronic cross sections in the TeV PeV regime with DAMPE through machine learning techniques.* PoS ICRC2023(these proceeding) 142