

Near-field radio emission induced by extensive air showers

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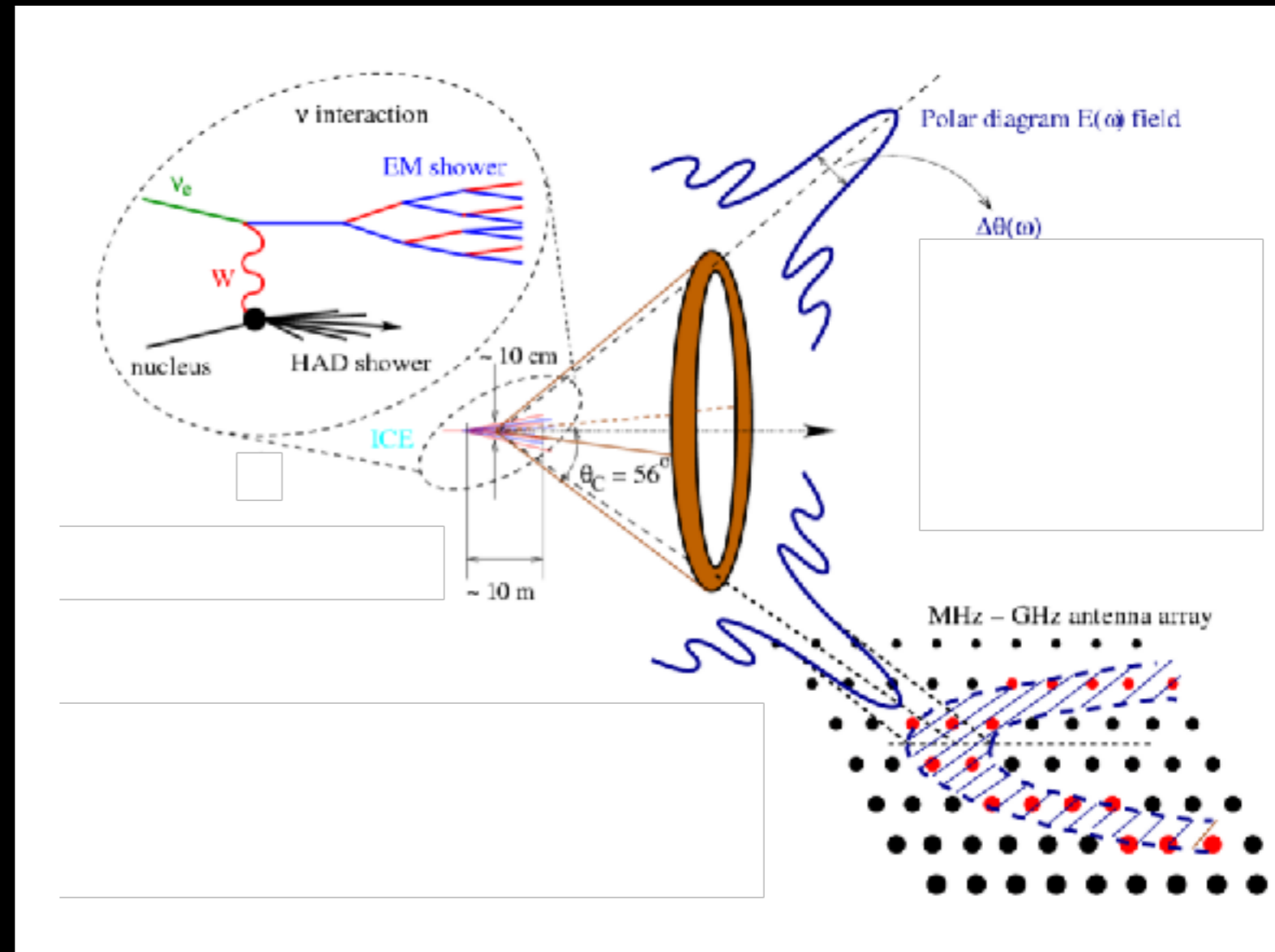


Take-home message

- **Low frequency** (< 10 MHz) radio emission of EAS needs a new treatment including **near-field** effects ($d \sim \lambda$)
- We expect the existence of a new signal called the **sudden death pulse** (SDP)
- We present a **formula suitable for** the calculation of the **low-frequency** electric field

Radio detection of cosmic rays (or neutrinos)

- A primary particle creates an EAS
- Charged particles in the EAS create electric field
- Electric field is measured (usually > 20 MHz)



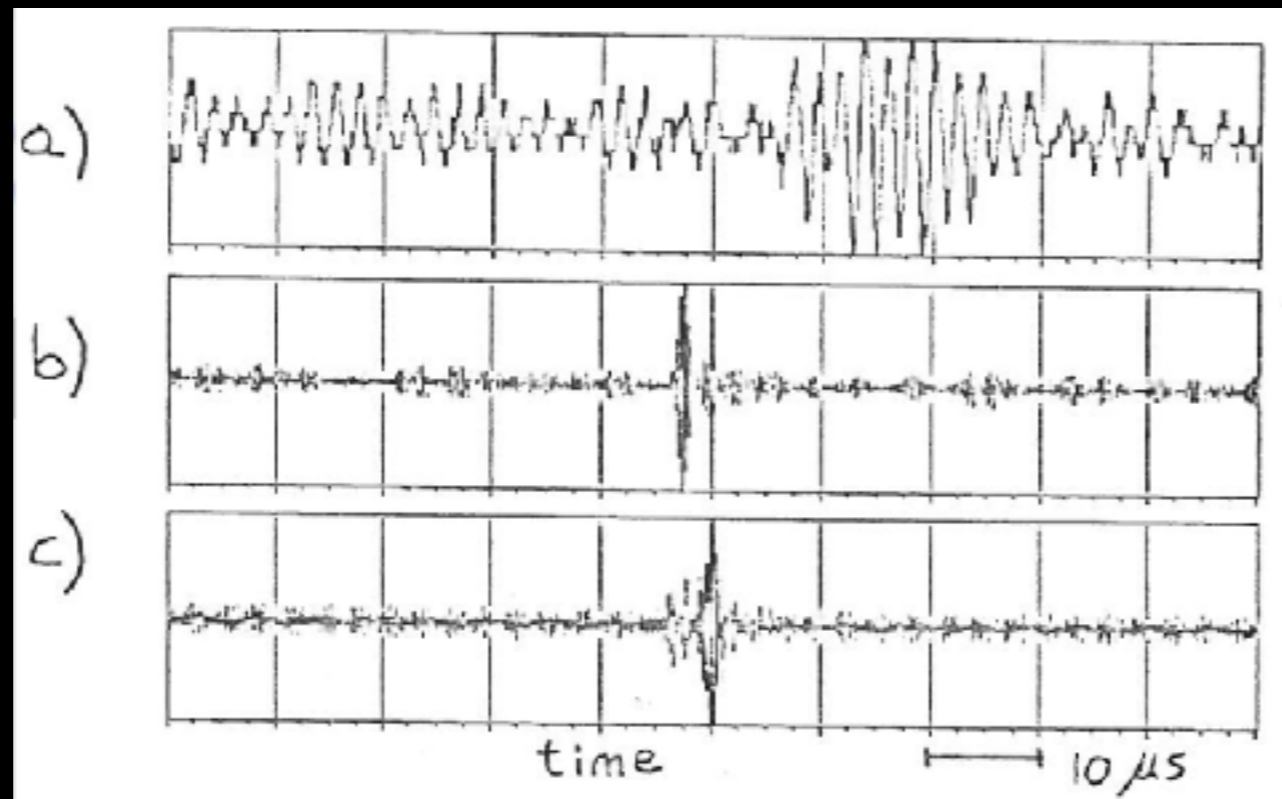
Why low frequency?

- Several experiments (EASTOP, Akeno) have **measured a large low-frequency emission** (C. Castagnoli et al., 22nd ICRC, 363. // K. Nishi, K. Suga. Proc. 20th ICRC (1987) 125)
- **Simulations** and **measurements** (see A. Escudie, [CRI102]) indicate an **emission** at low frequency with a **larger detection range**
- We expect a new kind of signal, the **sudden death pulse**

450 - 500 kHz

2.3 - 2.9 MHz

3.1 - 4.1 MHz



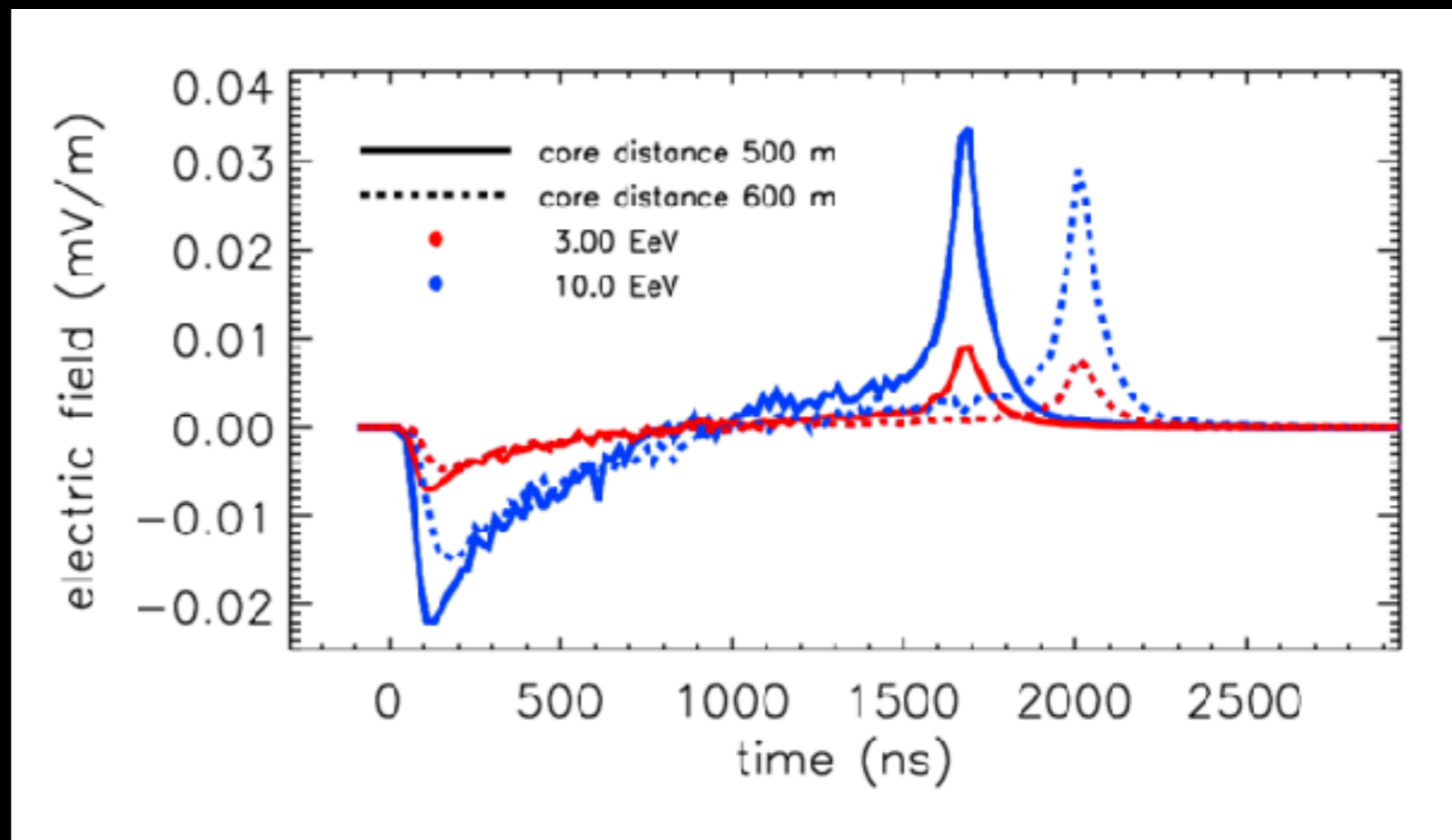
**EASRADIO (EASTOP)
Vertical polarisation**

Castagnoli et al., 22nd ICRC, 363

Sudden Death Pulse

- Shower particles are **decelerated** upon arrival to the **ground**
- Large shower footprint, but **coherence** at low frequencies (1 MHz ~ 300 m)
- Pulse at $t = d/c$ after shower core arrival
- Low-frequency pulse

Vertical component



ArXiv:1211.3305

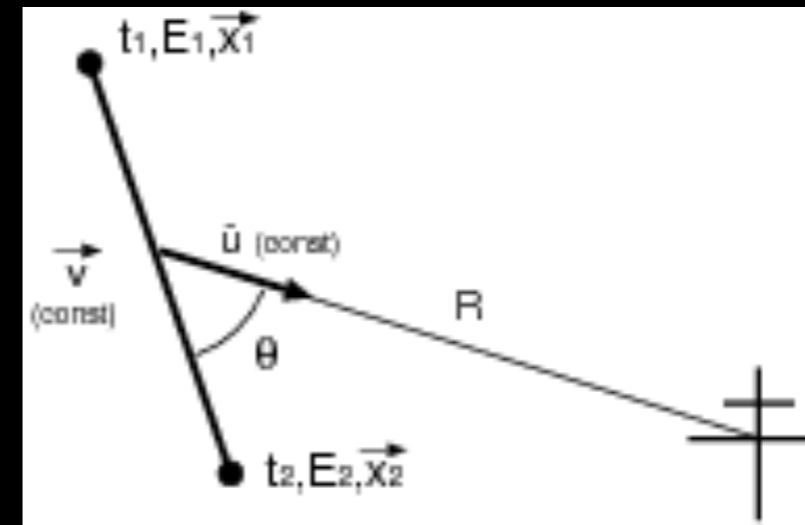
The EXTASIS experiment

- New experiment at the Nançay radio observatory
- **Detect** the **low-freq** (1.7 - 3.7 MHz) **counterpart** to the known EAS field
- Detect the **SDP**
- See [A. Escudie's talk](#) [CRI102]



Electric field for a track

- Codes such as SELFAS, ZHAireS or CoREAS use the **far-field** approximation ($kR \gg 1$)
- At 1 MHz, and $R = 100$ m: $kR \sim 2$.
Near field!
- Formula** for the field of a particle track at **all frequencies**:



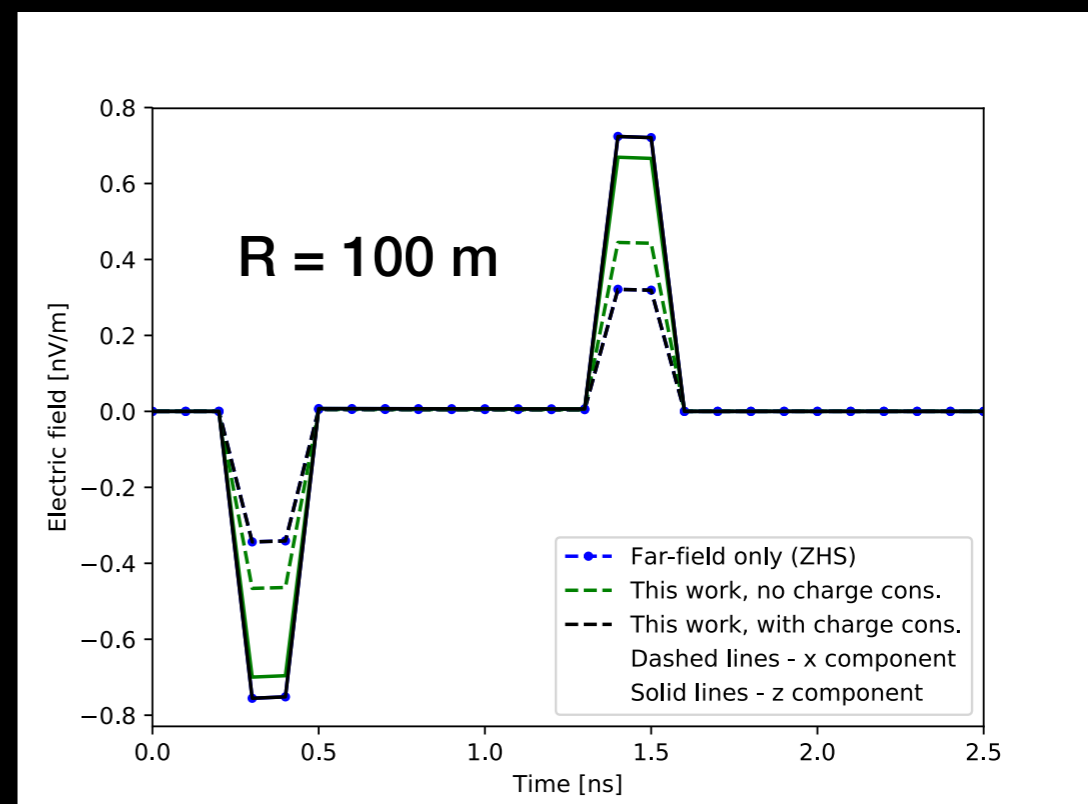
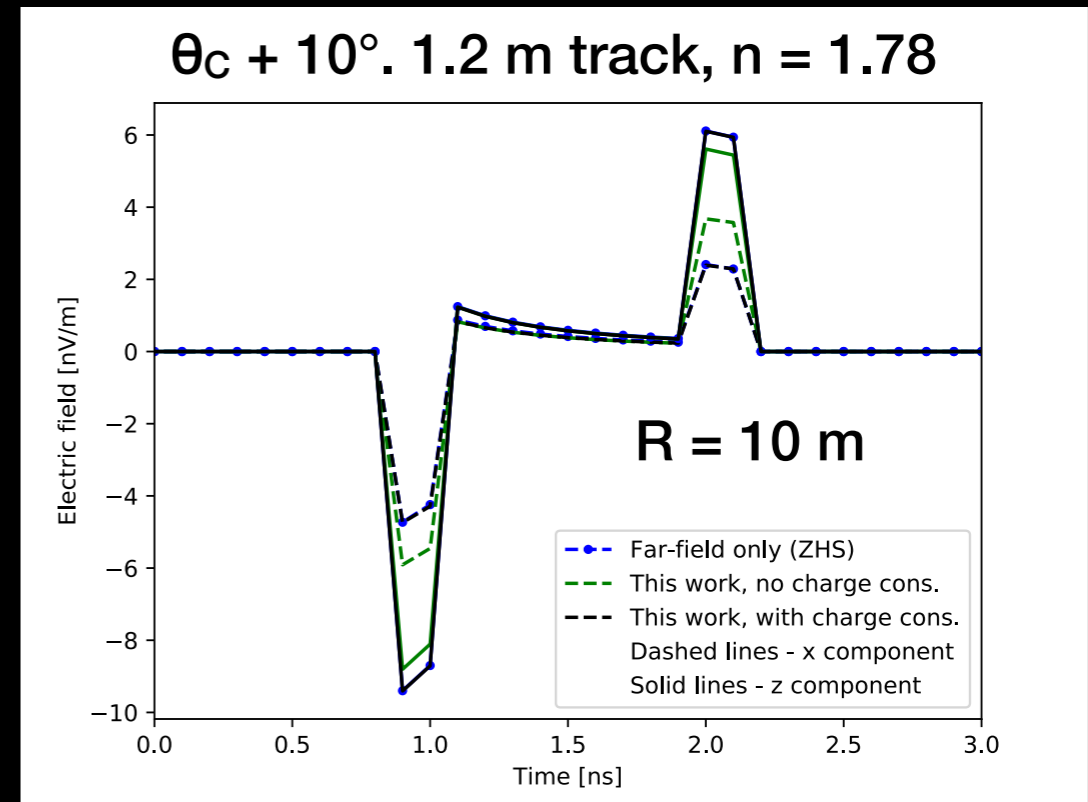
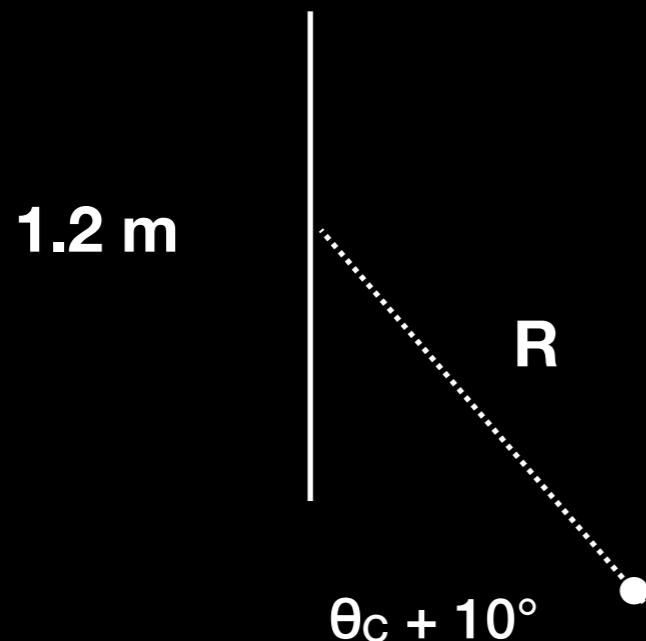
$$\mathbf{E}(\mathbf{x}, t) = \frac{1}{4\pi\epsilon} \int d^3x' \left\{ \left[\frac{\rho(\mathbf{x}', t_{\text{ret}})\mathbf{r}}{R^2(1 - n\boldsymbol{\beta} \cdot \mathbf{r})} \right]_{\text{ret}} + \frac{n}{c} \frac{\partial}{\partial t} \left[\frac{\rho(\mathbf{x}', t_{\text{ret}})\mathbf{r}}{R(1 - n\boldsymbol{\beta} \cdot \mathbf{r})} \right]_{\text{ret}} - \frac{n^2}{c^2} \frac{\partial}{\partial t} \left[\frac{\mathbf{J}(\mathbf{x}', t_{\text{ret}})}{R(1 - n\boldsymbol{\beta} \cdot \mathbf{r})} \right]_{\text{ret}} \right\}$$

- Caveat: **Charge MUST be conserved!**

$$\rho(\mathbf{x}, t) = q\delta^{(3)}(\mathbf{x} - \mathbf{x}_1)\Theta(t_1 - t) + q\delta^{(3)}(\mathbf{x} - \mathbf{x}_1 - \mathbf{v}(t - t_1))\Theta(t - t_1)\Theta(t_2 - t_1) + q\delta^{(3)}(\mathbf{x} - \mathbf{x}_2)\Theta(t - t_2)$$

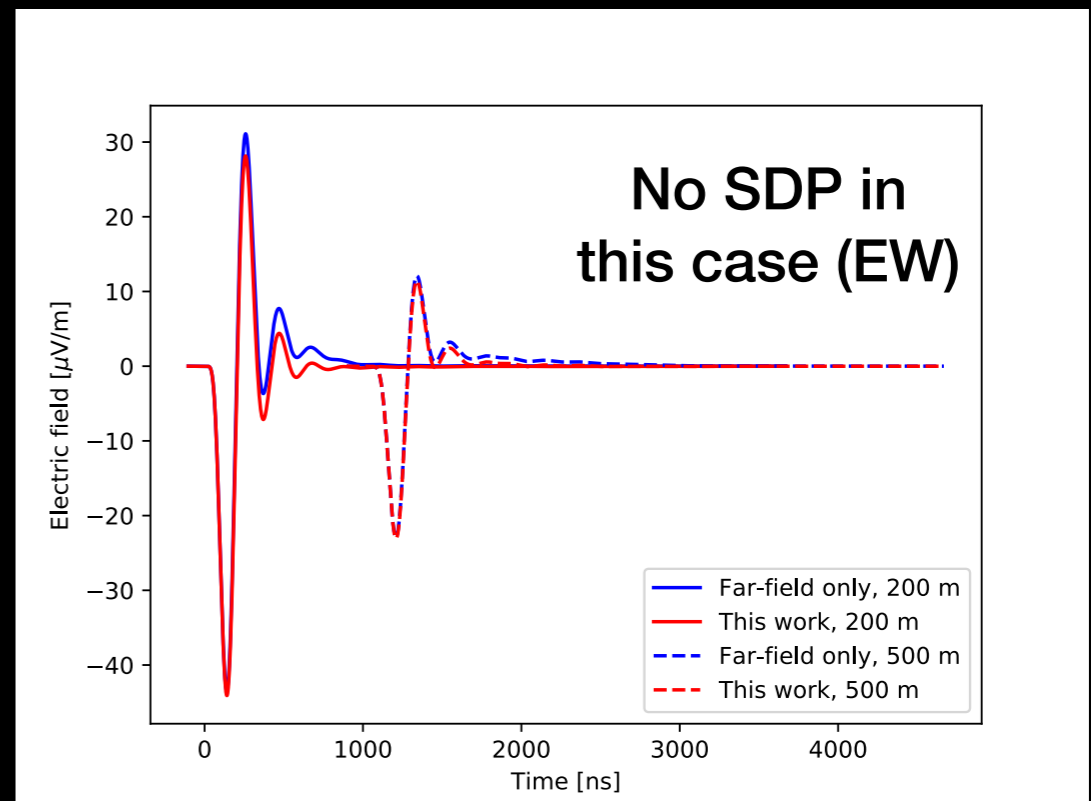
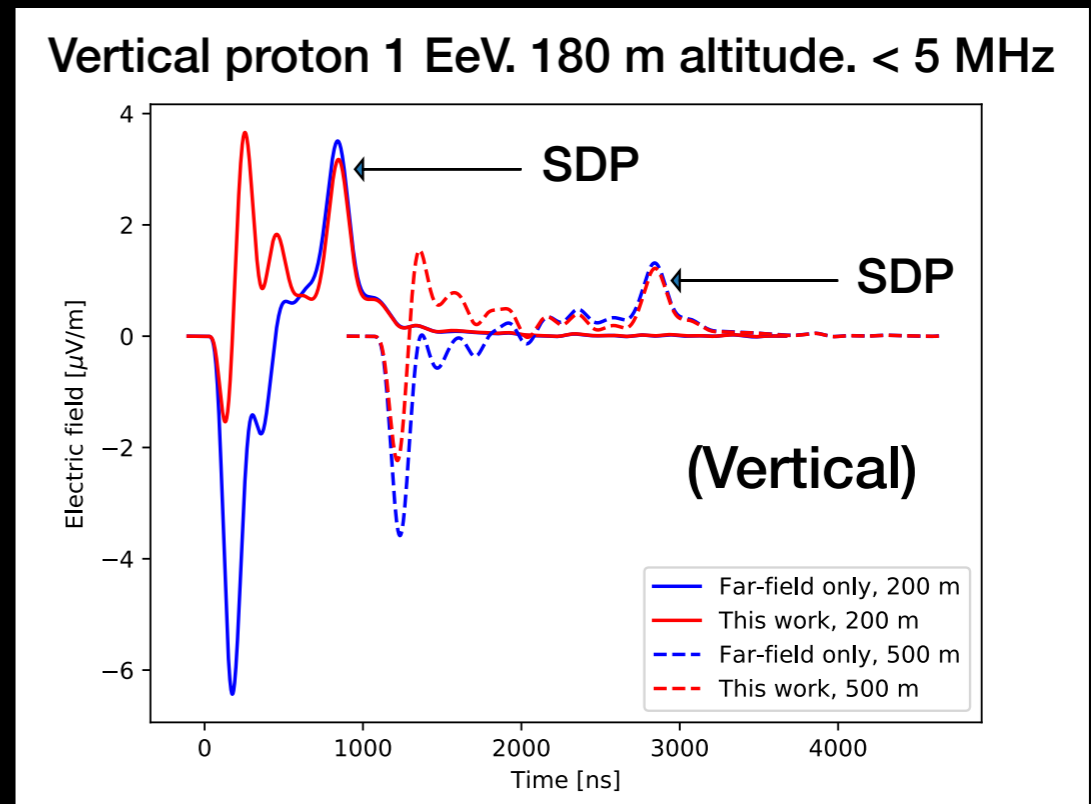
Comparison with far-field: ZHS

- Our formula yields the **same result** as the **ZHS** formula (far-field). J. Alvarez-Muniz et al. Phys. Rev. D 81 (2010) 123009
- If charge is not conserved, pulses from the beginning and end of the track are not well reproduced.



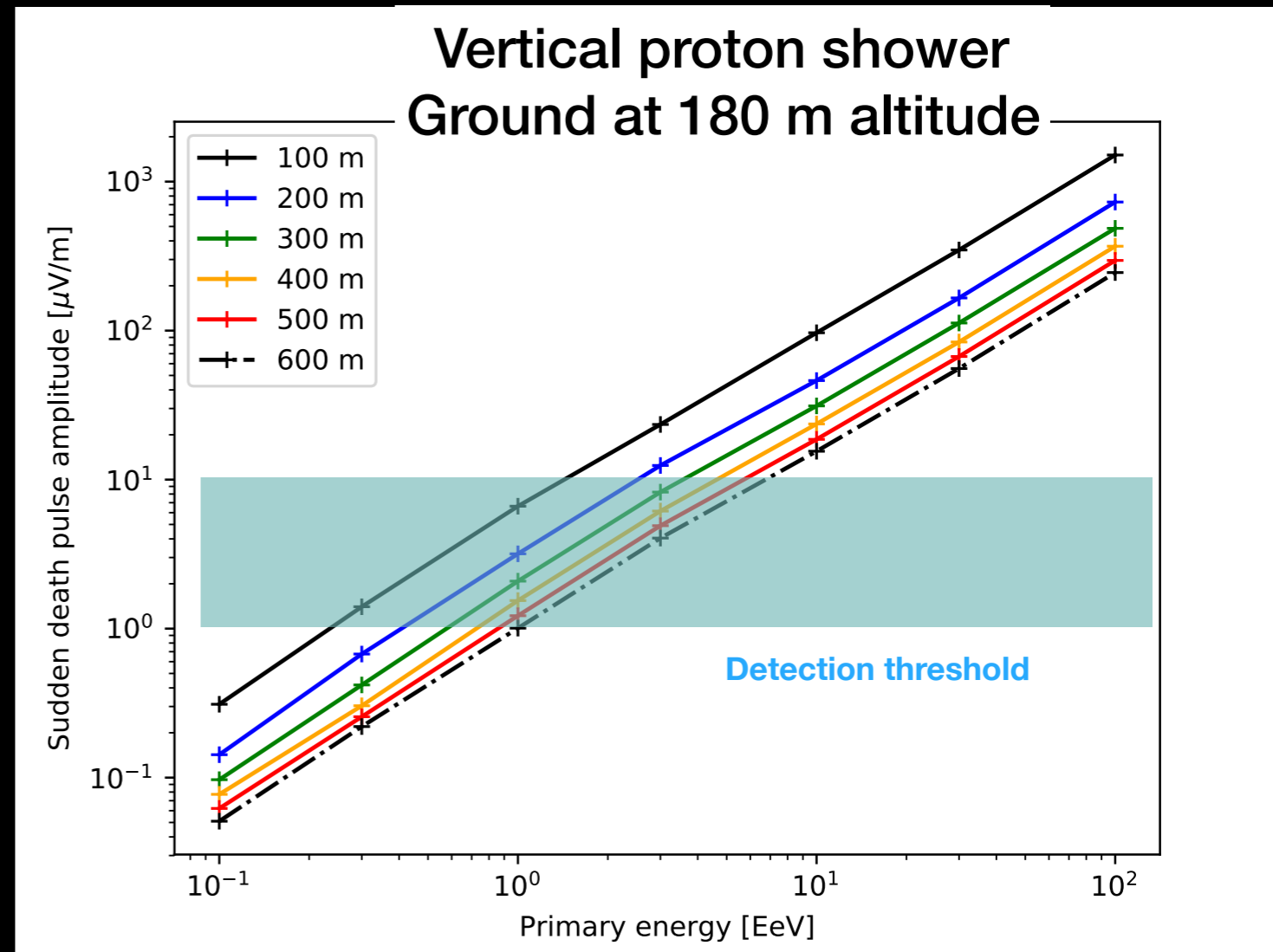
Implementation in SELFAS

- **SELFAS** (open source) is a **MC code** that calculates the **field of an EAS**. It has been upgraded with a state-of-the-art treatment of the atmosphere (see [B. Revenu \[CRI109\]](#))
- We have implemented our formula assuming:
 - **No static field** after shower extinction
 - Particles are suddenly **stopped at ground level**
 - **No reflection** (can be taken into account with antenna pattern) or **surface wave**
 - No transmission (attenuation in soil)



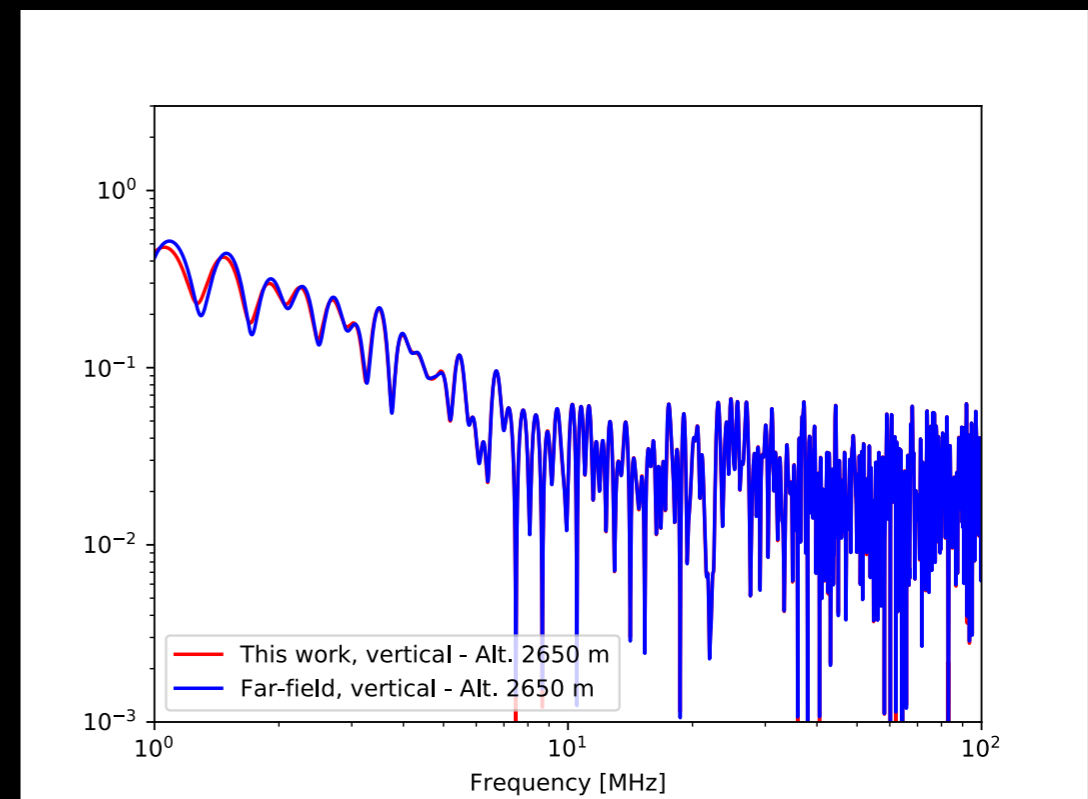
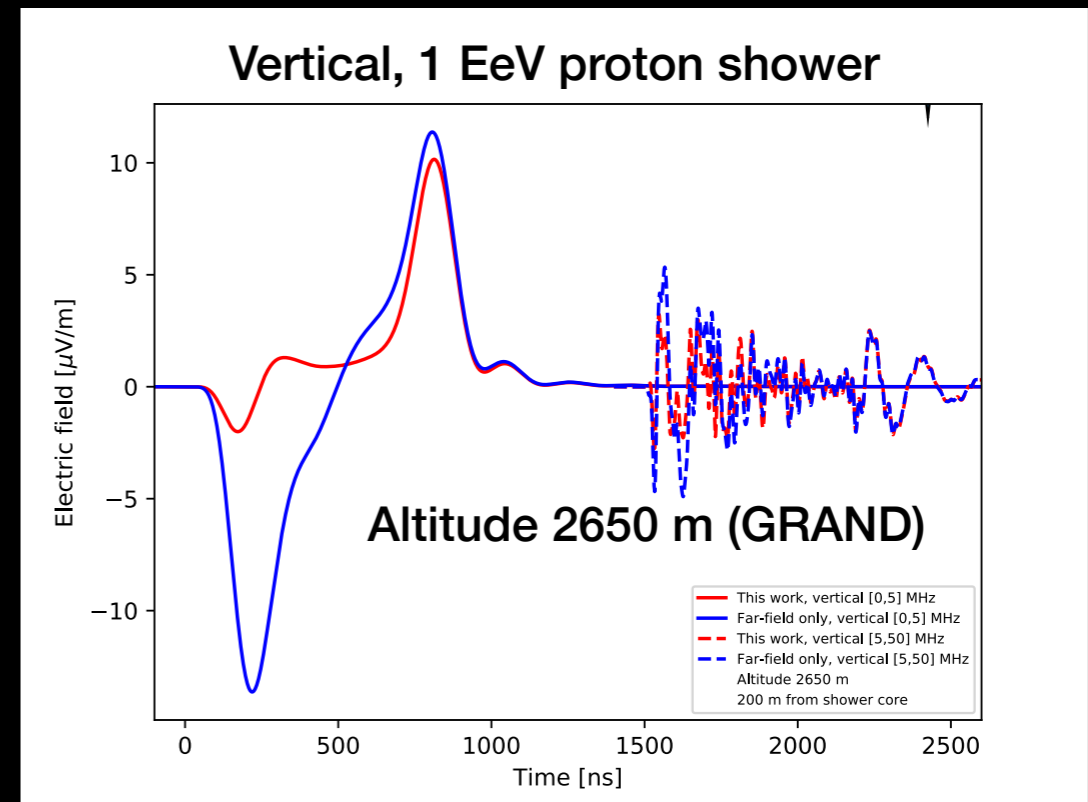
Sudden Death Pulse amplitude

- **SDP amplitude** calculated as a function of **energy and distance** (vertical proton shower)
- The amplitude is **proportional** to the **energy** (number of particles arriving to the ground)
- At Nançay, we expect detectability between 1 and 10 $\mu\text{V}/\text{m}$



Low frequency emission at high altitude

- The amplitude is proportional to the energy (number of particles arriving to the ground)
- **Altitude closer to the shower maximum** means **larger SDP** signal
- More total signal below 10 MHz!
See spectrum.



Conclusions

- We have presented an equation for the **electric field** of a particle track **valid for all frequencies** (includes near-field effects)
- Correctly taking into account near-field effects is **crucial for low frequency** measurements (below 10 MHz). That is the case for the **EXTASIS** experiment.
- We have **implemented** this formula in the **SELFAS** Monte Carlo code and checked its **consistency** with **far-field (ZHS)** calculations.
- An analysis of the surface wave and the effects of the interface on the field is underway.
- **Caveat:** we have talked about **electric field, not voltage**. The response of the antennas in the near field could be complicated; only far-field properties are usually well known.

Thank you
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Low frequency emission at high altitude

- The amplitude is proportional to the energy (number of particles arriving to the ground)
- **Higher** altitude means **larger SDP** signal
- More total signal below 10 MHz!
See spectrum.

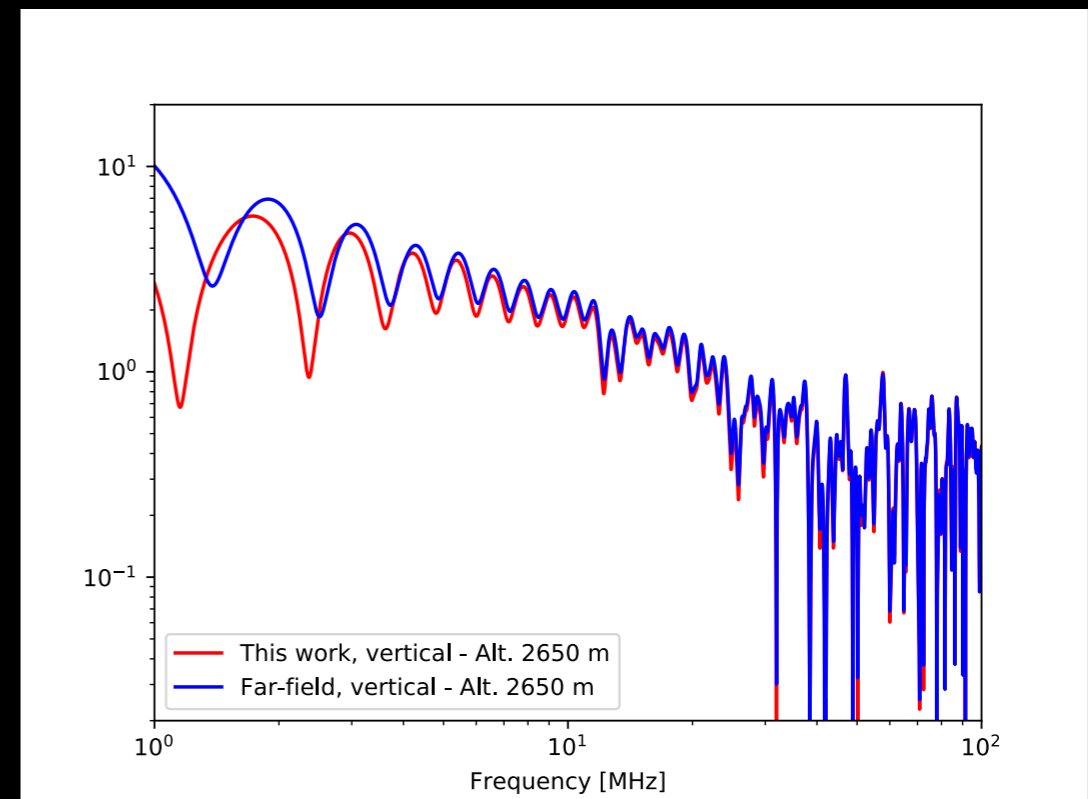
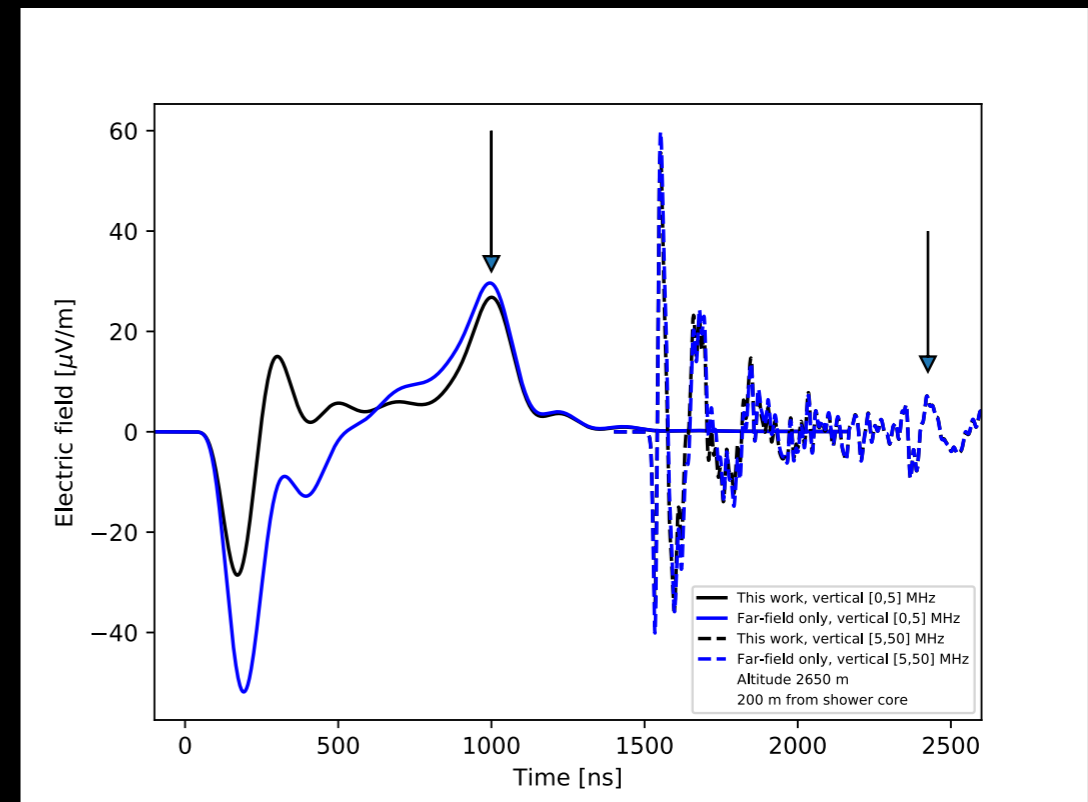


Figure: 30 degrees 5 EeV proton shower
Ground at 2650 m of altitude

Comparison with exact formula (frequency)

$$\mathbf{E}(\mathbf{x}, \omega) = \frac{q}{4\pi\epsilon} \left\{ \int_{t_1}^{t_2} dt e^{i\omega t} \frac{e^{ikR}}{R} \mathbf{r} \left[\frac{1}{R} - \frac{i\omega n}{c} \right] + \sum_{j=1,2} \left[(-1)^j e^{i\omega t_j} e^{ikR_j} \mathbf{r}_j \left(\frac{n}{cR_j} + \frac{i}{\omega R_j^2} \right) \right] \right\} + \frac{i\omega\mu_0 q}{4\pi} \int_{t_1}^{t_2} dt e^{i\omega t} \frac{e^{ikR}}{R} \mathbf{v}$$

- Our **formula** (above, in **frequency domain**) yields the **same result** as the exact formula in frequency domain in Phys. Rev D 87 (2013) 023003.
- Therefore, the formula in time domain reproduces the **complete electric field**

Figure: Cherenkov angle
1.2 m long track in ice, $n = 1.78$

