

Gauge Mediation in the NMSSM with a Light Singlet: Sparticles within the Reach of LHC Run II

Ben Allanach

DAMTP, CMS, Wilberforce Road, University of Cambridge, Cambridge, CB3 OWA, United Kingdom E-mail: b.c.allanach@damtp.cam.ac.uk

Marcin Badziak*

Institute of Theoretical Physics, Faculty of Physics, University of Warsaw ul. Pasteura 5, PL-02-093 Warsaw, Poland *E-mail:* Marcin.Badziak@fuw.edu.pl

Cyril Hugonie

LUPM, UMR 5299, CNRS, Université de Montpellier II, 34095 Montpellier, France E-mail: Cyril.Hugonie@univ-montp2.fr

Robert Ziegler

Sorbonne Universités, UPMC Univ Paris 06, UMR 7589, LPTHE, F-75005, Paris, France CNRS, UMR 7589, LPTHE, F-75005, Paris, France *E-mail:* robert.ziegler@lpthe.jussieu.fr

Relatively light stops in gauge mediation models are usually made compatible with the Higgs mass of 125 GeV by introducing direct Higgs-messenger couplings. We show that such couplings are not necessary in a simple and predictive model that combines minimal gauge mediation and the next-to-minimal supersymmetric standard model (NMSSM). We show that one can obtain a 125 GeV Standard Model-like Higgs boson with stops as light as 1.1 TeV, thanks to the mixing of the Higgs with a singlet state at $\mathcal{O}(90-100)$ GeV that can explain the LEP excess. In this scenario the singlet-higgs-higgs superfields coupling λ is small and tan β large. Sparticle searches at the LHC may come with additional b-jets or taus and may involve displaced vertices. The sparticle production cross-section at the 13 TeV LHC can be $\mathcal{O}(10-100)$ fb, leading to great prospects for discovery in the early phase of LHC Run II.

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*Speaker.

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1. Introduction

The discovery of a Higgs boson with mass of about 125 GeV [1] has considerable impact on supersymmetric (SUSY) model building. In its simplest realization, the Minimal Supersymmetric Standard Model (MSSM), the tree-level Higgs mass is bounded from above by the Z-boson mass, which implies that large radiative corrections of the order of the tree-level mass are needed [2]. In consequence, the predicted stop masses are generically in a multi-TeV range. In MSSM, the only way to keep the stop masses at the level of 1 TeV is to have large top trilinear coupling A_t , adjusted such that the stop-mixing contribution to the Higgs mass is maximal. However, large stop mixing cannot be present in gauge mediation models [3] because they predict negligible A-terms at the messenger scale. Large A_t at the messenger scale can be generated by introducing direct Higgs-messenger couplings [4]. This allows to accommodate the 125 GeV Higgs with relatively light stops but solving SUSY flavor problem in such a scenario requires additional model building, see e.g. Ref. [5].

It is interesting to ask whether it is possible to have relatively light stops with gauge mediated SUSY breaking without direct Higgs-messenger couplings. This might be possible only in extensions of MSSM with new contributions to the Higgs mass. In the context of gauge mediation, the Next-to-MSSM (NMSSM) [6] is particularly well motivated extension of MSSM because it provides a simple solution to the notorious $\mu - B_{\mu}$ problem [7]. Yet it is very difficult to realize this scenario with minimal gauge mediation (MGM), as the NMSSM soft terms are too small [8]. These problems can however be cured by adding direct couplings of the singlet to messengers, at the cost of a single new parameter without generating new sources of flavor violation (for a different approach see e.g. Ref. [9]). A viable model of this kind has been proposed by Delgado, Giudice and Slavich (DGS) in Ref. [10] (extensions with Higgs-messenger couplings were studied in Ref. [11]). However, the authors of Ref. [10] concluded that in this model sparticles cannot be lighter than in MGM.

In NMSSM, a possible source of enhancement of the Higgs mass is mixing with a singletdominated scalar that is lighter than the SM-like Higgs [12, 13]. In these proceedings, which are mainly based on the results of Ref. [14], we re-analyze the DGS model and identify new viable regions in the parameter space where singlet-Higgs mixing is small enough to pass experimental constraints, but large enough to give substantial contributions to the tree-level Higgs mass. This model can therefore rely on smaller contributions from stop loops, thus reducing the overall scale of sparticle masses. Interestingly, squarks and gluinos can be light enough to be discovered in the early stage of the LHC run II, in contrast to MGM, where a 125 GeV Higgs mass requires colored sparticles beyond the reach of the LHC (even for very high luminosity) [15]. Moreover, we find that the light singlet-like scalar can easily explain the 2σ excess around 98 GeV observed in the LEP Higgs searches [16, 17]. The realization of this scenario, with maximal contribution to the tree-level Higgs mass from mixing, fixes almost all of the model parameters. A single parameter remains free and controls the details of the phenomenology. The lightest supersymmetric particle (LSP) is the gravitino and the next-to-LSP (NLSP) is the singlino, a setup that leads to new signatures at collider experiments. The underlying model might therefore serve as a representative for a whole class of signatures that motivate suitable SUSY search strategies.

2. The DGS Model

The field content of the DGS model (see Ref. [10] for details) consists of the NMSSM fields (the MSSM fields plus a gauge singlet *S*), in addition to two copies of messengers in $\mathbf{5} + \mathbf{\bar{5}}$ of SU(5), denoted by $\Phi_i, \bar{\Phi}_i, i = 1, 2$ with SU(2) doublet and SU(3) triplet components $\Phi_i^D, \bar{\Phi}_i^T, \bar{\Phi}_i^T, i =$ 1,2. Supersymmetry breaking is parametrized by a non-dynamical background field $X = M + F\theta^2$. Apart from the Yukawa interactions, the superpotential is given by the NMSSM part, the spurionmessenger couplings and the singlet-messenger couplings, $W = W_{\text{NMSSM}} + W_{\text{GM}} + W_{\text{DGS}}$, where

$$W_{\rm NMSSM} = \lambda S H_u H_d + \frac{\kappa}{3} S^3 , \qquad (2.1)$$

$$W_{\rm GM} = X \sum_{i=1,2} \left(\kappa_i^D \bar{\Phi}_i^D \Phi_i^D + \kappa_i^T \bar{\Phi}_i^T \Phi_i^T \right), \qquad (2.2)$$

$$W_{\rm DGS} = S \left(\xi_D \bar{\Phi}_1^D \Phi_2^D + \xi_T \bar{\Phi}_1^T \Phi_2^T \right) \,. \tag{2.3}$$

The new couplings in Eq. (2.3) are assumed to unify at the GUT scale $\xi_D(M_{\text{GUT}}) = \xi_T(M_{\text{GUT}}) \equiv \xi$. A similar assumption can be made for $\kappa_i^{D,T}$, but these parameters are largely irrelevant for the spectrum.

 \mathscr{Z}_3 invariant NMSSM models such as this one have a potential cosmological problem with domain walls, which are predicted to appear during the phase transition associated with electroweak symmetry breaking. This can be solved by the introduction of higher dimensional \mathscr{Z}_3 -violating operators. Dangerous tadpoles that may threaten successful electroweak symmetry breaking may be avoided by the imposition of a discrete \mathscr{R} -symmetry on the higher dimension operators [18].

Through the superpotential in Eq. (2.2) the messengers feel SUSY breaking at tree-level and communicate it to the NMSSM fields via gauge interactions and the direct couplings in Eq. (2.3). The contribution from gauge interactions is given by the usual expressions of MGM for one-loop gaugino masses M_i and two-loop sfermion masses $m_{\tilde{f}}$ at the messenger scale M

$$M_i = 2g_i^2 \tilde{m}, \qquad \qquad m_{\tilde{f}}^2 = 4\sum_{i=1}^3 C_i(f) \ g_i^4 \tilde{m}^2, \qquad (2.4)$$

where $\tilde{m} \equiv 1/(16\pi^2)F/M$ and $C_i(f)$ is the quadratic Casimir of the representation of the field f under SU(3) × SU(2) × U(1). The contributions from direct singlet-messenger couplings generate one-loop A-terms for the NMSSM couplings

$$A_{\lambda} = \frac{A_{\kappa}}{3} = -\tilde{m}\left(2\xi_D^2 + 3\xi_T^2\right), \qquad (2.5)$$

and two-loop contributions to soft masses for the singlet and the Higgs fields

$$\begin{split} \tilde{m}_{S}^{2} &= \tilde{m}^{2} \left[8\xi_{D}^{4} + 15\xi_{T}^{4} + 12\xi_{D}^{2}\xi_{T}^{2} \right] \\ &- \tilde{m}^{2} \left[4\kappa^{2} \left(2\xi_{D}^{2} + 3\xi_{T}^{2} \right) \right] \\ &- \tilde{m}^{2} \left[\xi_{D}^{2} \left(\frac{6}{5}g_{1}^{2} + 6g_{2}^{2} \right) + \xi_{T}^{2} \left(\frac{4}{5}g_{1}^{2} + 16g_{3}^{2} \right) \right], \end{split}$$

$$(2.6)$$

$$\Delta \tilde{m}_{H_u}^2 = \Delta \tilde{m}_{H_d}^2 = -\tilde{m}^2 \lambda^2 \left(2\xi_D^2 + 3\xi_T^2 \right) \,. \tag{2.7}$$

Marcin Badziak

There is also a one-loop contribution to the singlet soft mass [10] that is relevant only for very low messenger scales and has negligible impact on the spectrum. The model is thus determined by five parameters: \tilde{m} , M, λ , κ and ξ , where one parameter (following DGS we choose κ) can be eliminated by requiring correct electroweak symmetry breaking (EWSB).

3. Low-energy Spectrum

There are several regions in the parameter space that lead to a viable spectrum and are compatible with perturbative couplings up to the GUT scale. They can be distinguished by the size of the relative contributions to the SM-like Higgs mass, which are given schematically by

$$m_h^2 = M_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta + m_{h,\text{mix}}^2 + m_{h,\text{loop}}^2, \qquad (3.1)$$

where $m_{h,\text{mix}}^2$ is the contribution from mixing with the other two CP-even states in the full Higgs mass matrix.

Since a larger tree-level mass implies a lighter SUSY spectrum, we concentrate here on the region in parameter space where the effective tree-level contribution to the Higgs mass is maximized. It turns out that one can reach up to $m_h^2 - m_{h,loop}^2 \approx (99 \,\text{GeV})^2$, provided that $\tan \beta$ is large and the contribution from mixing is sizable and positive. This requires the singlet state to be lighter than the SM-like Higgs, which is not excluded by LEP if the mixing angle with the SM-Higgs is small enough. Note that the LHC constrains this scenario only through measurements of the SM-like Higgs couplings that are suppressed by the mixing.

A focal feature of this scenario is its high level of predictivity, as three out of four free parameters of the model are determined by the Higgs sector. Maximizing the tree-level contribution to the SM-like Higgs mass m_{h_2} fixes the singlet-like Higgs mass m_{h_1} and the singlet-Higgs mixing angle θ (mixing with the heavy Higgs doublet is negligible) along the lines of Ref. [13], giving approximately $m_{h_1} \approx 94 \text{ GeV}$ and $\cos \theta \approx 0.88$. This in turn determines the model parameters λ and ξ , while the overall scale of soft terms \tilde{m} is fixed by the required size of the residual loop contribution to m_{h_2} . Departing from the maximal tree-level contribution leads to slightly different m_{h_1} and θ , and of course larger \tilde{m} .

In practice one can map the model parameters to the low-energy spectrum only numerically. For this analysis we have used a modified version of NMSSMTools [19]. Independent checks using a modified version of SOFTSUSY3.4.1 [20] produced Higgs and SUSY spectra that agreed at the percent level.

Before discussing our results, we provide some rough analytic results that can be obtained neglecting renormalization group (RG) effects and expanding the NMSSM vacuum conditions in the limit of large singlet vacuum expectation value. In this way one obtains approximate relations between ξ , λ and the physical Higgs parameters

$$\xi \sim \frac{m_{h_1}}{4\sqrt{2}g_3\tilde{m}}, \qquad \qquad \lambda \sim \frac{m_{h_2}^2 - m_{h_1}^2}{4v\tilde{m}}\sin 2\theta. \qquad (3.2)$$

For TeV-scale superpartners (and the above values for m_{h_1} and θ that maximize the Higgs tree-level contribution) one finds $\xi \sim \lambda \sim 10^{-2}$.

The smallness of ξ and λ has important consequences for the low-energy spectrum. Small ξ implies small values of $|A_{\lambda}|$ and $|A_{\kappa}|$ and imposing proper EWSB yields $\kappa \ll \lambda$ and the prediction of large tan $\beta \sim \lambda/\kappa$. In turn, the smallness of A_{κ} and κ results in a very light singlet-like pseudoscalar a_1 of mass

$$m_{a_1} \sim \sqrt{\frac{45\sqrt{8\xi}}{32g_3}} m_{h_1}.$$
 (3.3)

For $\xi \sim 10^{-2}$, Eq. (3.3) predicts m_{a_1} to be smaller than m_{h_1} by a factor of a few. We find numerically that the light pseudoscalar mass varies between 20 and 40 GeV, for gluino masses below 2.5 TeV.

For the above range of parameters the mass of the singlino can be obtained from the following approximate sum rule [21]:

$$m_{\tilde{S}}^2 \approx m_{h_1}^2 + \frac{1}{3}m_{a_1}^2,$$
 (3.4)

which implies that the singlino mass is about 100 GeV. Since in MGM the LSP is the gravitino and the typical scale of the NLSP is the bino mass $M_1 \approx 420 \text{ GeV} (\tilde{m}/\text{ TeV})$, it is clear that here the singlino strongly dominates the composition of the NLSP. This is a distinguishing feature of this model.

This is closely connected to the main virtue of this scenario, the large contribution to the tree-level Higgs mass from singlet-Higgs mixing. This requires smaller radiative corrections from stop loops, and in turn much lighter sparticle masses than in MGM. Through these corrections the observed Higgs mass of 125 GeV essentially fixes the overall scale of the sparticle spectrum \tilde{m} , up to an estimated theoretical uncertainty of 3 GeV in the prediction of m_{h_2} . We find that $m_{h_2} = 125$



Figure 1: (Left panel) SM-like Higgs mass vs. gluino mass and (Right panel) singlet-like Higgs mass vs. lightest stop mass. The various model points are distinguished by the Higgs-singlet mixing angle θ , which decreases from top to bottom as specified in the left panel. For the same SM-like Higgs mass a larger mixing angle allows for much lighter gluinos. The lightest stop masses are obtained for a singlet-like Higgs around 94 GeV.

GeV is compatible with a gluino mass of 2.1 TeV (1.4 TeV if the theoretical uncertainty of 3 GeV for the Higgs mass is taken into account). Squarks of the first two generations have approximately the same mass as the gluino and should be within the reach of the LHC Run II. Stop masses can be as light as 1.7 (1.1) TeV, which should be compared with the lower bound on stop masses of about 8 (3) TeV in MGM [15]. In Fig. 1 we show the values of Higgs, gluino and stop masses for several model points separated by the singlet-Higgs mixing angle θ . Note that $\cos^2 \theta$ is roughly of the size

of effective Higgs signal strengths $R_i = (\sigma \times BR)_i / (\sigma \times BR)_i^{SM}$, which are substantially reduced in this scenario. Nevertheless all shown points are compatible with LEP and LHC constraints on the Higgs sector.

Having fixed $(\xi, \lambda, \tilde{m})$ by the set of physical Higgs parameters $(m_{h_1}, m_{h_2}, \theta)$, the only free parameter left is the messenger scale M. This parameter controls the low-energy spectrum in several ways. First of all, increasing M leads to larger values of A_t at the electroweak (EW) scale, which (as in MGM) is purely radiatively generated and therefore grows with the length of the RG running. In turn, this enhances the stop-mixing contribution to the Higgs mass, and therefore larger M leads to lighter stops and hence smaller \tilde{m} for the same value of m_{h_2} . Also, the value of M essentially determines the nature of the next-to-NLSP (NNLSP). For small $M \leq 10^8$ GeV the (mostly right-handed) stau is the NNLSP (with selectron and smuon being co-NNLSP), because the soft mass $m_{\tilde{E}}$ is smaller than M_1 at the messenger scale. For $M \gtrsim 10^9$ GeV (requiring gluino masses below 2.5 TeV) the RG effects are strong enough to raise $m_{\tilde{E}}$ above M_1 and the bino-like neutralino becomes the NNLSP. In the transition region 10^8 GeV $\lesssim M \lesssim 10^9$ GeV the NNLSP can be either stau or bino, depending on the other parameters. The messenger scale M controls the gravitino mass according to:

$$m_{3/2} = 38 \,\mathrm{eV}\left(\frac{\tilde{m}}{\mathrm{TeV}}\right) \left(\frac{M}{10^6 \,\mathrm{GeV}}\right).$$
 (3.5)

The simultaneous presence of gravitino LSP and singlino NLSP leads to a novel phenomenology quite different both from MGM models and from typical NMSSM scenarios.

4. LHC Phenomenology

In our scenario decay chains of every supersymmetric particle produced at the LHC end up in a singlino-like neutralino \tilde{N}_1 . Since the singlet couples very weakly, these decays always proceed through the NNLSP or co-NNLSP. The singlino subsequently decays to the gravitino and the singlet-like pseudoscalar a_1 , which in turn predominantly decays to *b*-quarks:

$$\tilde{N}_1 \to a_1 \tilde{G} \to b \bar{b} \tilde{G}$$
. (4.1)

The decay length of the neutralino (in its rest frame) is approximately given by

$$c\tau_{\tilde{N}_1} \approx 2.5 \,\mathrm{cm} \left(\frac{100 \,\mathrm{GeV}}{M_{\tilde{N}_1}}\right)^5 \left(\frac{M}{10^6 \,\mathrm{GeV}}\right)^2 \left(\frac{\tilde{m}}{\mathrm{TeV}}\right)^2.$$
 (4.2)

Since *M* cannot be much below 10^6 GeV, it is clear from the above formula that the singlino NLSP (with mass about 100 GeV) always travels macroscopic distance before it decays. For large *M* the singlino decays well outside the detector so it is stable from the collider point of view. However, for $M \sim 10^6 - 10^7$ GeV the singlino may decay in the detector after traveling some distance from the interaction point leading to a displaced vertex. Since the value of *M* also decides about the nature of the NNLSP, it can be used to define three regions with distinct LHC phenomenology, which we briefly discuss in the remainder of this article. A more detailed analysis of LHC phenomenology and discovery prospects will be the subject of a future publication [22].

| | P1 | P2 | P3 | P4 | P5 |
|--|---------------------|---------------------|--------------------|---------------------|---------------------|
| ĩ | $7.5 \cdot 10^2$ | $8.7 \cdot 10^{2}$ | $9.3 \cdot 10^2$ | $5.9 \cdot 10^{2}$ | $9.3 \cdot 10^{2}$ |
| М | $1.4 \cdot 10^{6}$ | $2.8\cdot 10^6$ | $3.3\cdot10^7$ | $8.3\cdot10^{14}$ | $3.4\cdot10^{14}$ |
| λ | $1.0 \cdot 10^{-2}$ | $9.3 \cdot 10^{-3}$ | $6.7\cdot10^{-3}$ | $9.2 \cdot 10^{-3}$ | $6.9\cdot10^{-3}$ |
| ξ | $1.2 \cdot 10^{-2}$ | $1.1\cdot 10^{-2}$ | $1.3\cdot 10^{-2}$ | $3.2\cdot10^{-2}$ | $2.0\cdot 10^{-2}$ |
| $\tan\beta$ | 25 | 28 | 24 | 26 | 21 |
| m_{h_1} | 92 | 93 | 98 | 94 | 94 |
| m_{h_2} | 122.1 | 123.4 | 122.9 | 122.1 | 125.0 |
| m_{a_1} | 26 | 26 | 28 | 40 | 32 |
| $m_{	ilde{N}_1}$ | 101 | 102 | 106 | 104 | 104 |
| $m_{\tilde{N}_2}$ | 322 | 377 | 400 | 251 | 379 |
| $m_{\tilde{e}_1}$ | 303 | 358 | 406 | 449 | 676 |
| $m_{	ilde{	au}_1}$ | 284 | 333 | 376 | 432 | 637 |
| $m_{	ilde{g}}$ | 1.73 | 1.98 | 2.09 | 1.37 | 2.06 |
| $m_{\tilde{u}_R}$ | 1.79 | 2.06 | 2.15 | 1.36 | 2.07 |
| $m_{\tilde{t}_1}$ | 1.64 | 1.87 | 1.90 | 1.06 | 1.63 |
| $c \tau_{\tilde{N}_1}$ | $6.4 \cdot 10^{-2}$ | 0.34 | 48 | $1.9 \cdot 10^{16}$ | $6.0 \cdot 10^{15}$ |
| $\sigma_{\tilde{q}\tilde{q}}^{13\mathrm{TeV}}$ | 9.35 | 2.99 | 1.98 | 59.7 | 2.63 |
| $\sigma_{\widetilde{q}\widetilde{g}}^{\widehat{1}3\mathrm{TeV}}$ | 11.9 | 3.30 | 2.01 | 91.1 | 2.48 |
| $\sigma_{\rm strong}^{13{ m TeV}}$ | 25.2 | 7.28 | 4.58 | 190 | 5.95 |
| $\sigma_{ m strong}^{ m 8TeV}$ | 0.51 | 0.07 | 0.03 | 10.1 | 0.05 |
| $\sigma_{ m EW}^{ m 13TeV}$ | 27 | 12 | 7.5 | 6.7 | 5.6 |
| $\sigma_{ m EW}^{ m 8TeV}$ | 5.5 | 2.1 | 1.2 | 1.3 | 0.7 |

Table 1: List of benchmark points. All masses are in GeV except colored sparticle masses in TeV, the neutralino decay length $c\tau_{\tilde{N}_1}$ in m and cross-sections in fb. All points have reduced effective Higgs couplings, with Higgs signal strengths about 0.75, as a result of a Higgs-singlet mixing angle with $\cos \theta \approx 0.88$.

In Table 1 we collect several characteristic benchmark points. Points P1 and P4 represent the lightest SUSY spectra we have found, for very low and very large messenger scales, respectively. Since the Higgs mass errors are pushed to the limits, we consider these points merely as limiting cases, although not necessarily unrealistic. Note in particular that P4 is not obviously ruled out by standard SUSY searches for jets + missing E_T , since the additional decay of the would-be-LSP bino to singlino reduces efficiency compared to the CMSSM [21, 23]. The other points are representatives for the three characteristic regions discussed below, and P3 is in addition chosen to fit the LEP excess. Note that all points have quite large singlet-Higgs mixing, leading to reduced effective Higgs couplings. Points with smaller mixing and/or larger Higgs masses can be obtained by increasing the overall SUSY scale \tilde{m} .

In all regions sparticles can be very light, so that huge parts of the parameter space are in the reach of LHC Run II. As can be seen from Table 1 the total strong production cross-section (dominated by $\tilde{q}\tilde{q}$ and $\tilde{q}\tilde{g}$) is $\mathcal{O}(10-100)$ fb, as computed with PROSPINO [24]. LHC Run II is expected to deliver $\mathcal{O}(10)$ fb⁻¹ of integrated luminosity in 2015, which results in $\mathcal{O}(100-1000)$ potentially discoverable events. The total EW production cross-section at the 13 TeV LHC (computed with Pythia 8.2 [25]) is typically comparable to the strong one but is distributed among many different channels with rather small individual cross-sections of order $\mathcal{O}(1-10)$ fb. The most frequent EW production channel is $\chi_1^+ \chi_3^0$ (which are wino-like states decaying dominantly to staus) with the cross-section of about one fifth of the total EW cross-section ¹.

4.1 Low-M Region: $M \lesssim 10^7 \, \text{GeV}$

In this region, represented by benchmarks P1 and P2 in Table 1, the lightest stau is the NNLSP (with smuon/selectron co-NNLSPs) and therefore the singlino is produced in association with either tau or leptons. Since the splitting between sleptons and the singlino is around 200 GeV or more, one expects high- p_T taus or leptons in the final state, which presumably can be used to reduce QCD backgrounds considerably. In this region the singlino decays (via light pseudoscalar) to $b\bar{b}$ still inside the detector. However, identifying these displaced *b*-jets might be challenging since they are expected to be very soft due to the small pseudoscalar mass. We note that the low-M region is constrained (as are all GMSB models) by the matter power spectrum as inferred from the Lyman- α forest data and WMAP [26], which disfavours a gravitino mass between 16 eV and $m_{3/2}^{crit} \sim \mathcal{O}(\text{keV})$. In fact, from our scan we do not find any gravitinos with mass less than 16 eV. The other bound implies that

$$c\tau_{\tilde{N}_1} \gtrsim 17 \,\mathrm{m} \left(\frac{100 \,\mathrm{GeV}}{M_{\tilde{N}_1}}\right)^5 \left(\frac{m_{3/2}^{\mathrm{crit}}}{\mathrm{keV}}\right)^2$$
(4.3)

from Eqs. (3.5),(4.2). We remind the reader that $M_{\tilde{N}_1}$ is close to 100 GeV because of the sum rule Eq. (3.4). P1 and P2 violate this bound, whereas P3 may or may not, depending on the precise value of $m_{3/2}^{\text{crit}}$. However, entropy production after gravitino decoupling may provide a cosmological evasion of the bound for any point [27, 28].

4.2 Medium-M Region: $10^7 \text{ GeV} \lesssim M \lesssim 10^9 \text{ GeV}$

In this region, represented by benchmark P3, the singlino LSP is long-lived. Stau is still NNLSP, but smuon/selectron are no longer co-NNLSPs because they are heavier than the bino-like neutralino. In consequence, a vast majority of gluino and squark decay chains ends in stau NNLSP decaying to tau and quasi-stable singlino NLSP, with two high- p_T taus in each event.

4.3 Large-M Region: $M \gtrsim 10^9 \,\text{GeV}$

For large messenger scales, represented by benchmarks P4 and P5, the NNLSP is bino-like. Therefore the (quasi-stable) singlino is typically produced in association with the 125 GeV Higgs, $BR(\tilde{N}_2 \rightarrow \tilde{N}_1 h_2) \sim 70 - 75\%$, or the singlet-like Higgs, $BR(\tilde{N}_2 \rightarrow \tilde{N}_1 h_1) \sim 25 - 30\%$. Both Higgs states decay dominantly to $b\bar{b}$. Using a b-jet tagging efficiency of 70% [29], one still expects

¹At the 8 TeV LHC the EW production cross-section is larger than the strong production cross-section (except for P4 where it is comparable). Nevertheless, current LHC limits for direct EW production are far too weak to constrain the model. The dominant EW production channel is a production of wino-like charginos and neutralinos with masses of about 600-700 GeV (except for P4), which subsequently decay dominantly to staus. The lower mass limits for charginos decaying into staus has been set in some simplified models but are always below 400 GeV [30, 31].

in each event at least two (three) identified high- p_T b-jets from bino decays with a probability of about 60 (30)%. This comes on top of the b-jets originating from other decays in the gluino and/or squark decay chains. Therefore, this model can be easily discriminated against MSSM models using searches with large numbers of b-jets.

5. Conclusions

In this article we have re-analyzed the DGS model, a simple and predictive framework for combining MGM and the NMSSM. We have found new regions in the parameter space with a singlet at $\mathcal{O}(90-100)$ GeV and singlet-Higgs mixing giving substantial contributions to the tree-level Higgs mass. While these regions are compatible with Higgs precision data, we find that colored sparticles can be close to their direct search limits, in sharp contrast to MGM.

SUSY decays have more visible particles and can lead to less missing E_T as compared to MSSM predictions. The phenomenology is controlled by a single parameter, which determines whether SUSY decays chains lead to additional b-jets or taus and involve displaced vertices. Particularly interesting signatures of this model are displaced vertices originating from a singlino-like neutralino NLSP decay to gravitino LSP and a very light pseudoscalar Higgs (predominantly decaying to $b\bar{b}$). As the production cross-sections of colored sparticles are $\mathcal{O}(10-100)$ fb, significant parts of parameter space are discoverable in the early stage of LHC Run II.

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References

- ATLAS Collaboration, Phys. Lett. B 716 (2012) 1 [arXiv:1207.7214 [hep-ex]]; CMS Collaboration, Phys. Lett. B 716 (2012) 30 [arXiv:1207.7235 [hep-ex]].
- [2] L. J. Hall, D. Pinner and J. T. Ruderman, JHEP 1204 (2012) 131 [arXiv:1112.2703 [hep-ph]].
- [3] For a review see G. F. Giudice and R. Rattazzi, Phys. Rept. 322 (1999) 419 [hep-ph/9801271].
- [4] J. L. Evans, M. Ibe and T. T. Yanagida, Phys. Lett. B 705, 342 (2014) [arXiv:1107.3006 [hep-ph]].
- [5] L. Calibbi, P. Paradisi and R. Ziegler, JHEP 1306, 052 (2013) [arXiv:1304.1453 [hep-ph]].
- [6] For a review see U. Ellwanger, C. Hugonie and A. M. Teixeira, Phys. Rept. 496 (2010) 1 [arXiv:0910.1785 [hep-ph]].
- [7] G. R. Dvali, G. F. Giudice and A. Pomarol, Nucl. Phys. B 478 (1996) 31 [hep-ph/9603238].

- [8] M. Dine and A. E. Nelson, Phys. Rev. D 48 (1993) 1277 [hep-ph/9303230]; A. de Gouvea,
 A. Friedland and H. Murayama, Phys. Rev. D 57 (1998) 5676 [hep-ph/9711264].
- [9] V. S. Mummidi and S. K. Vempati, Nucl. Phys. B 881 (2014) 181 [arXiv:1311.4280 [hep-ph]].
- [10] A. Delgado, G. F. Giudice and P. Slavich, Phys. Lett. B 653 (2007) 424 [arXiv:0706.3873 [hep-ph]].
- [11] N. Craig, S. Knapen, D. Shih and Y. Zhao, JHEP 1303 (2013) 154 [arXiv:1206.4086 [hep-ph]].
- [12] See e.g. K. Agashe, Y. Cui and R. Franceschini, JHEP 1302 (2013) 031 [arXiv:1209.2115 [hep-ph]];
 U. Ellwanger and C. Hugonie, JHEP 1408 (2014) 046 [arXiv:1405.6647 [hep-ph]].
- [13] M. Badziak, M. Olechowski and S. Pokorski, JHEP 1306 (2013) 043 [arXiv:1304.5437 [hep-ph]].
- [14] B. Allanach, M. Badziak, C. Hugonie and R. Ziegler, Phys. Rev. D 92, no. 1, 015006 (2015) [arXiv:1502.05836 [hep-ph]].
- [15] P. Draper, P. Meade, M. Reece and D. Shih, Phys. Rev. D 85 (2012) 095007 [arXiv:1112.3068 [hep-ph]]; M. A. Ajaib, I. Gogoladze, F. Nasir and Q. Shafi, Phys. Lett. B 713 (2012) 462 [arXiv:1204.2856 [hep-ph]].
- [16] ALEPH and DELPHI and L3 and OPAL Collaborations and LEP Working Group for Higgs Boson Searches, Eur. Phys. J. C 47 (2006) 547 [arXiv:hep-ex/0602042]
- [17] G. Belanger, U. Ellwanger, J. F. Gunion, Y. Jiang, S. Kraml and J. H. Schwarz, JHEP 1301 (2013) 069 [arXiv:1210.1976 [hep-ph]].
- [18] C. Panagiotakopoulos and K. Tamvakis, Phys. Lett. B 446, 224 (1999) [hep-ph/9809475].
- [19] U. Ellwanger, J. F. Gunion and C. Hugonie, JHEP 0502 (2005) 066 [hep-ph/0406215]; U. Ellwanger and C. Hugonie, Comput. Phys. Commun. 175 (2006) 290 [hep-ph/0508022].
- [20] B. C. Allanach, Comput. Phys. Commun. 143, 305 (2002) [hep-ph/0104145]; B. C. Allanach,
 P. Athron, L. C. Tunstall, A. Voigt and A. G. Williams, Comput. Phys. Commun. 185, 2322 (2014)
 [arXiv:1311.7659 [hep-ph]].
- [21] D. Das, U. Ellwanger and A. M. Teixeira, JHEP 1204 (2012) 067 [arXiv:1202.5244 [hep-ph]].
- [22] B. Allanach, M. Badziak, G. Cottin, N. Desai, C. Hugonie and R. Ziegler, in preparation.
- [23] U. Ellwanger and A. M. Teixeira, JHEP 1410 (2014) 113 [arXiv:1406.7221 [hep-ph]].
- [24] W. Beenakker, R. Hopker, M. Spira and P. M. Zerwas, Nucl. Phys. B 492 (1997) 51
 [hep-ph/9610490]; W. Beenakker, M. Kramer, T. Plehn, M. Spira and P. M. Zerwas, Nucl. Phys. B 515 (1998) 3 [hep-ph/9710451].
- [25] T. Sjostrand, S. Mrenna and P. Z. Skands, JHEP 0605 (2006) 026 [hep-ph/0603175]; T. SjÂŽstrand, S. Ask, J. R. Christiansen, R. Corke, N. Desai, P. Ilten, S. Mrenna and S. Prestel *et al.*, Comput. Phys. Commun. 191 (2015) 159 [arXiv:1410.3012 [hep-ph]].
- [26] M. Viel, J. Lesgourgues, M. G. Haehnelt, S. Matarrese and A. Riotto, Phys. Rev. D 71, 063534 (2005) [astro-ph/0501562].
- [27] E. A. Baltz and H. Murayama, JHEP 0305, 067 (2003) [astro-ph/0108172].
- [28] M. Fujii and T. Yanagida, Phys. Lett. B 549, 273 (2002) [hep-ph/0208191].
- [29] CMS Collaboration, JINST 8 (2013) P04013 [arXiv:1211.4462 [hep-ex]].
- [30] G. Aad et al. [ATLAS Collaboration], JHEP 1410 (2014) 96 [arXiv:1407.0350 [hep-ex]].
- [31] V. Khachatryan *et al.* [CMS Collaboration], Eur. Phys. J. C 74 (2014) 9, 3036 [arXiv:1405.7570 [hep-ex]].