

# ESO VLT optical spectroscopy of BL Lac objects

**Boris Sbarufatti**

*INAF, Istituto di Astrofisica Spaziale e Fisica Cosmica di Palermo, Via Ugo La Malfa 153,  
I-90146 Palermo, Italy*

*E-mail:* sbarufatti@ifc.inaf.it

**Stefano Ciprini**

*1. Physics Department University of Perugia, and I.N.F.N. Perugia Section, Via A. Pascoli,  
I-06123 Perugia, Italy*

*2. Tuorla Observatory, University of Turku, Väisäläntie 20, FIN-21500 Piikkiö, Finland*

*E-mail:* stefano.ciprini@pg.infn.it

**Jari Kotilainen**

*Tuorla Observatory, University of Turku, Väisäläntie 20, FIN-21500 Piikkiö, Finland*

*E-mail:* jarkot@utu.fi

**Roberto Decarli**

*Università degli Studi dell'Insubria, via Valleggio 11, I-22100 Como, Italy*

*E-mail:* roberto.decarli@mib.infn.it

**Aldo Treves\***

*Università degli Studi dell'Insubria, via Valleggio 11, I-22100 Como, Italy*

*E-mail:* aldo.treves@uninsubria.it

**Angelo Veronesi**

*Università degli Studi dell'Insubria, via Valleggio 11, I-22100 Como, Italy*

*E-mail:* averonesi@dfm.uninsubria.it

**Renato Falomo**

*INAF, Osservatorio Astronomico di Padova, Vicolo dell'Osservatorio 5, I-35122 Padova, Italy*

*E-mail:* renato.falomo@oapd.inaf.it

Using ESO VLT plus FORS 1 for spectroscopy of BL Lac objects of unknown  $z$ , we measured the redshift of 4 targets and lower limits for 8 others.

*Workshop on Blazar Variability across the Electromagnetic Spectrum*

*April 22-25 2008*

*Palaiseau, France*

---

\*Speaker.

## Observations, analysis and results

We present results of an ongoing program at the ESO VLT plus FORS 1 for high S/N spectroscopy of BL Lac objects lacking of a firm redshift estimate, performed in non optimal seeing conditions (see Sbarufatti et al. 2005<sup>[1]</sup>, 2006<sup>[2]</sup>).

We report on 12 new sources for which we confirm the BLL classification. New redshifts are determined for 4 objects, 2 with weak emission lines (PKS 1057-79,  $z = 0.569$ ; TXS 2346+052,  $z = 0.419$ ) and 2 with absorptions from the host galaxy (RBS 1752,  $z = 0.449$ ; RBS 1915,  $z = 0.243$ ), see Table 1 and Figure 1.

For the remaining 8 BL Lacs, from the very absence of absorption lines of the host galaxy, lower limits to the redshift are deduced with  $z_{min}$  in the interval 0.20 - 0.80, see Sbarufatti et al. 2006<sup>[2]</sup> and Table 2.

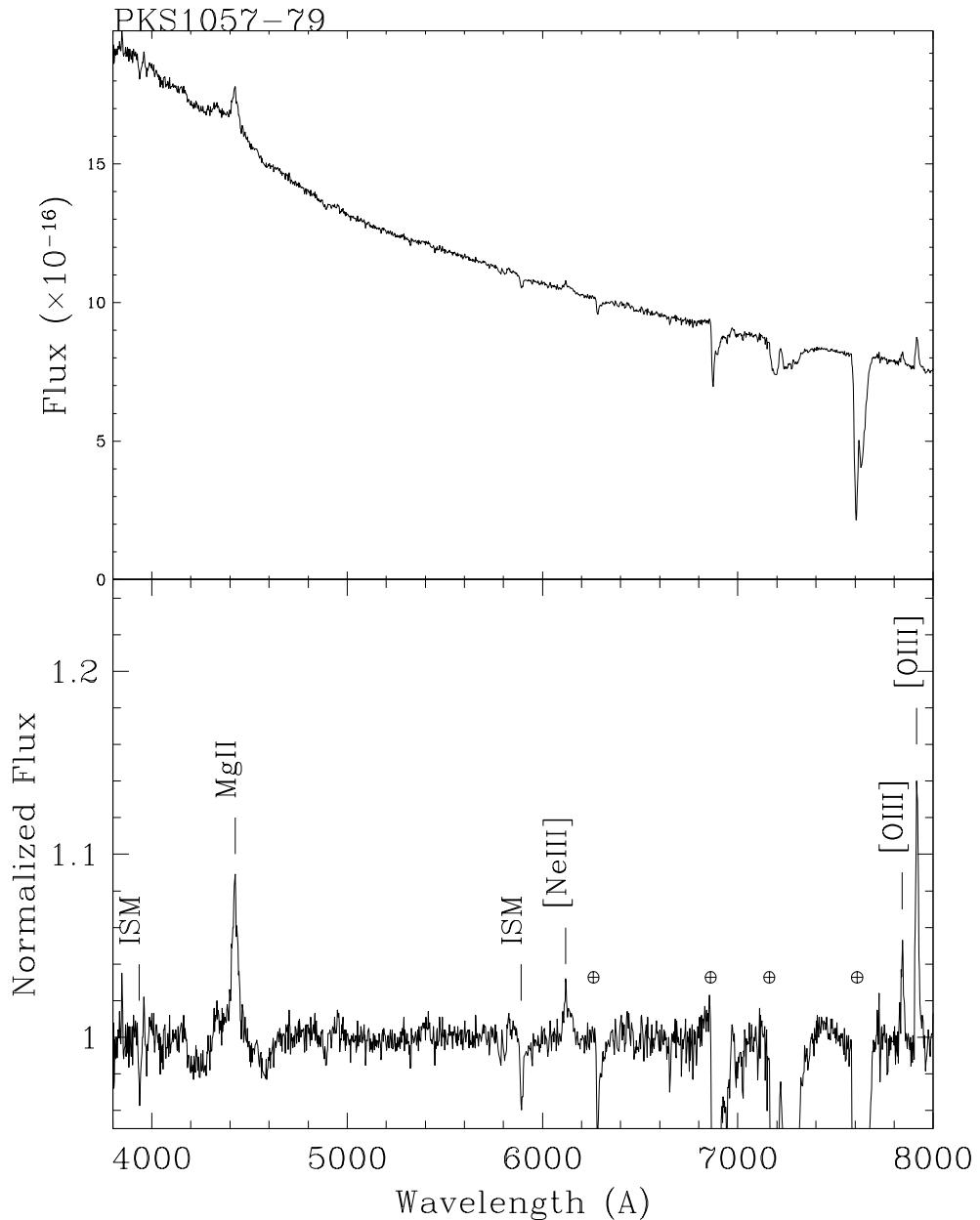
A detailed description of the adopted techniques and results is contained in Sbarufatti et al. 2008<sup>[3]</sup>. All the spectra of our program can be retrieved at <http://www.oapd.inaf.it/zblac/>.

Object name	redshift
PKS 1057-79	$z = 0.569$
RBS 1752	$z = 0.449$
RBS 1915	$z = 0.243$
TXS 2346+052	$z = 0.419$

**Table 1:** Redshifts for BL Lac objects

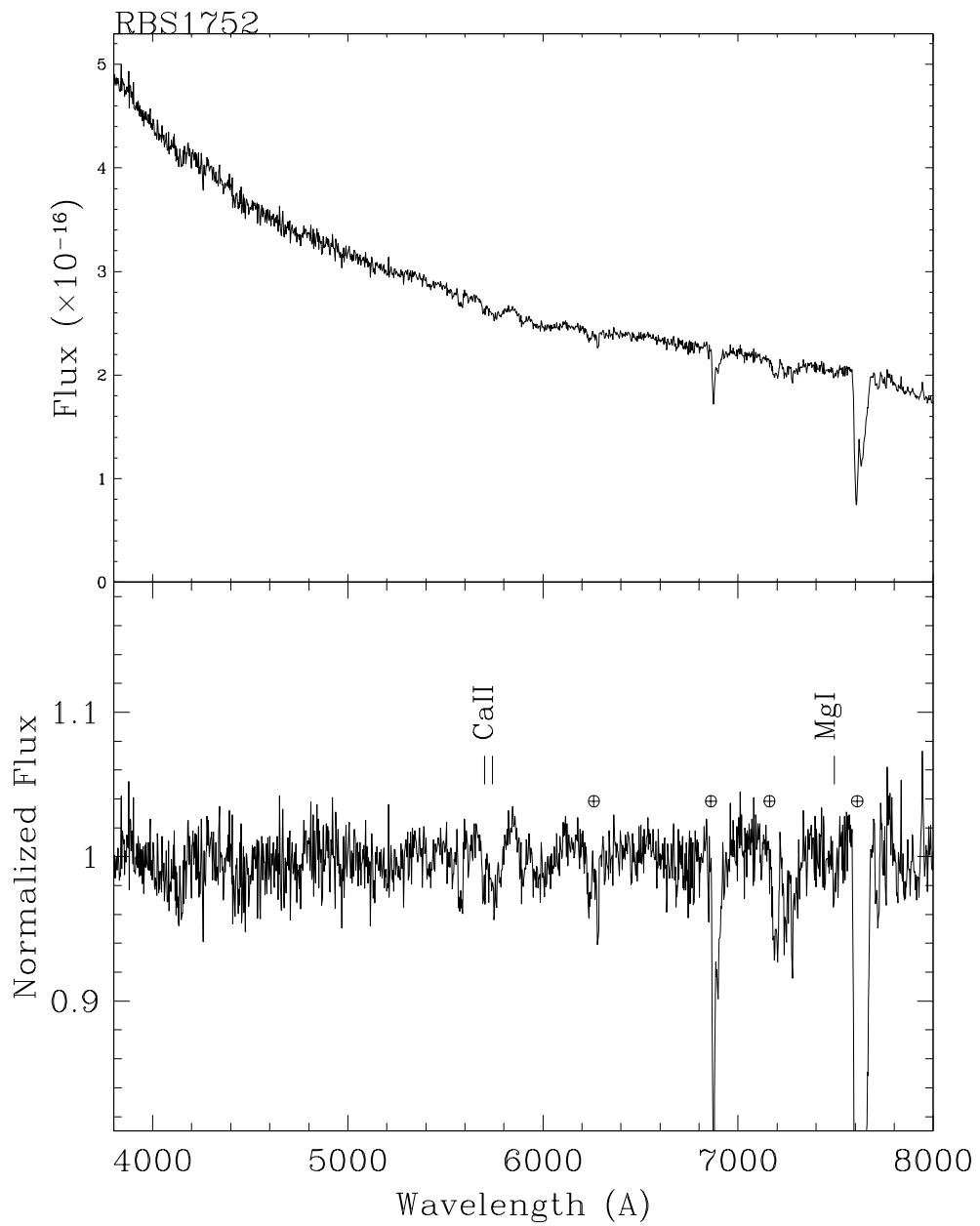
Object name	lower limits to the redshift
PKS 0019+058	$z > 0.4$
GC 0109+224	$z > 0.2$
RBS 0231	$z > 0.4$
OM 280	$z > 0.2$
OQ 012	$z > 0.5$
PMNJ 1539–0658	$z > 0.8$
PKS 1830–589	$z > 0.5$
1RXS J235730.1–171801	$z > 0.6$

**Table 2:** Lower limits to the redshifts for BL Lac objects without spectral features

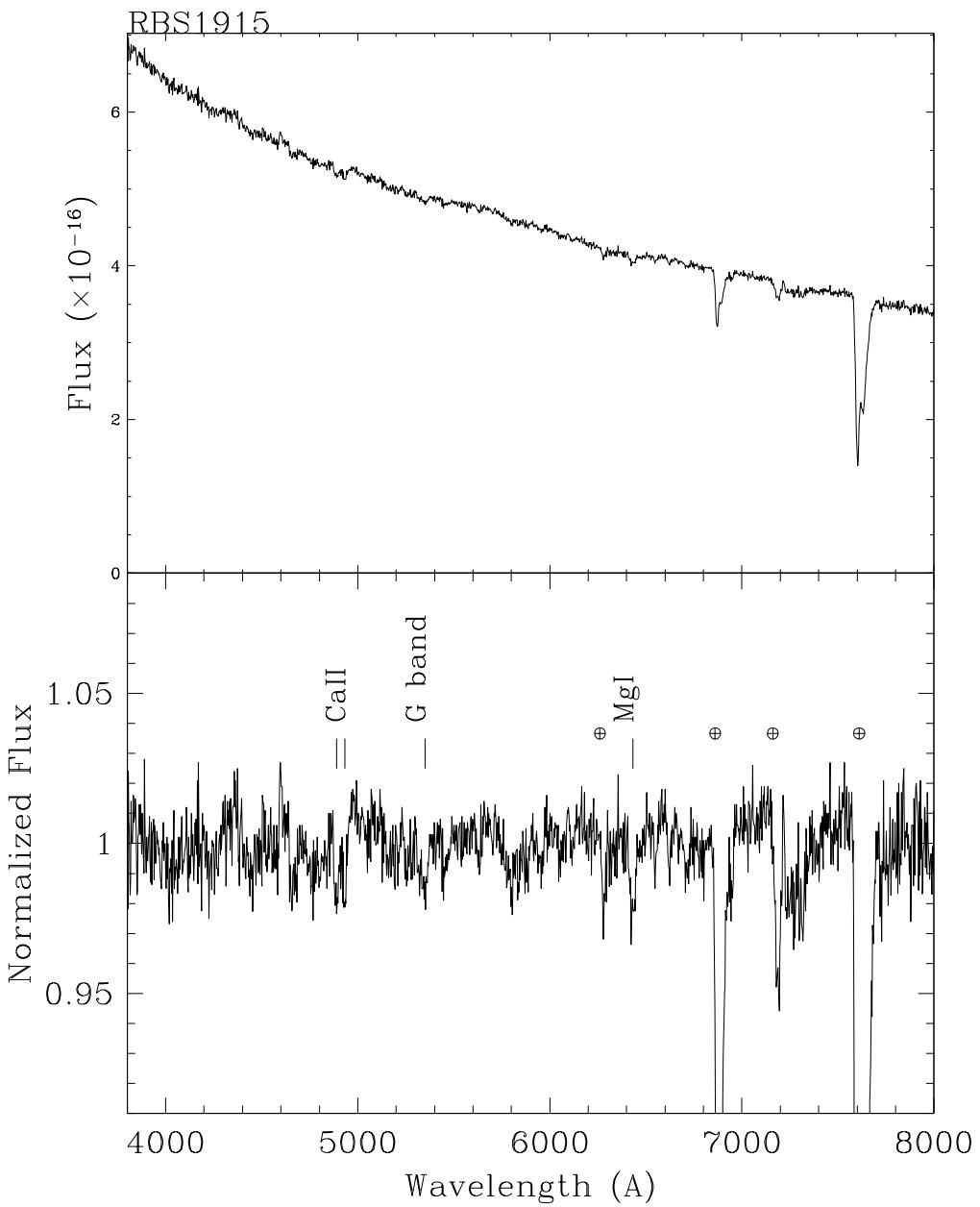


**Figure 1:** Top panels: flux calibrated dereddened spectra; the flux is measured in units of  $\text{erg}/\text{cm}^2/\text{s}/\text{\AA}$ . Bottom panels: normalized spectra. Telluric bands are indicated by  $\oplus$ , spectral features are marked by the line identifications.

PoS (BLAZARS2008) 026

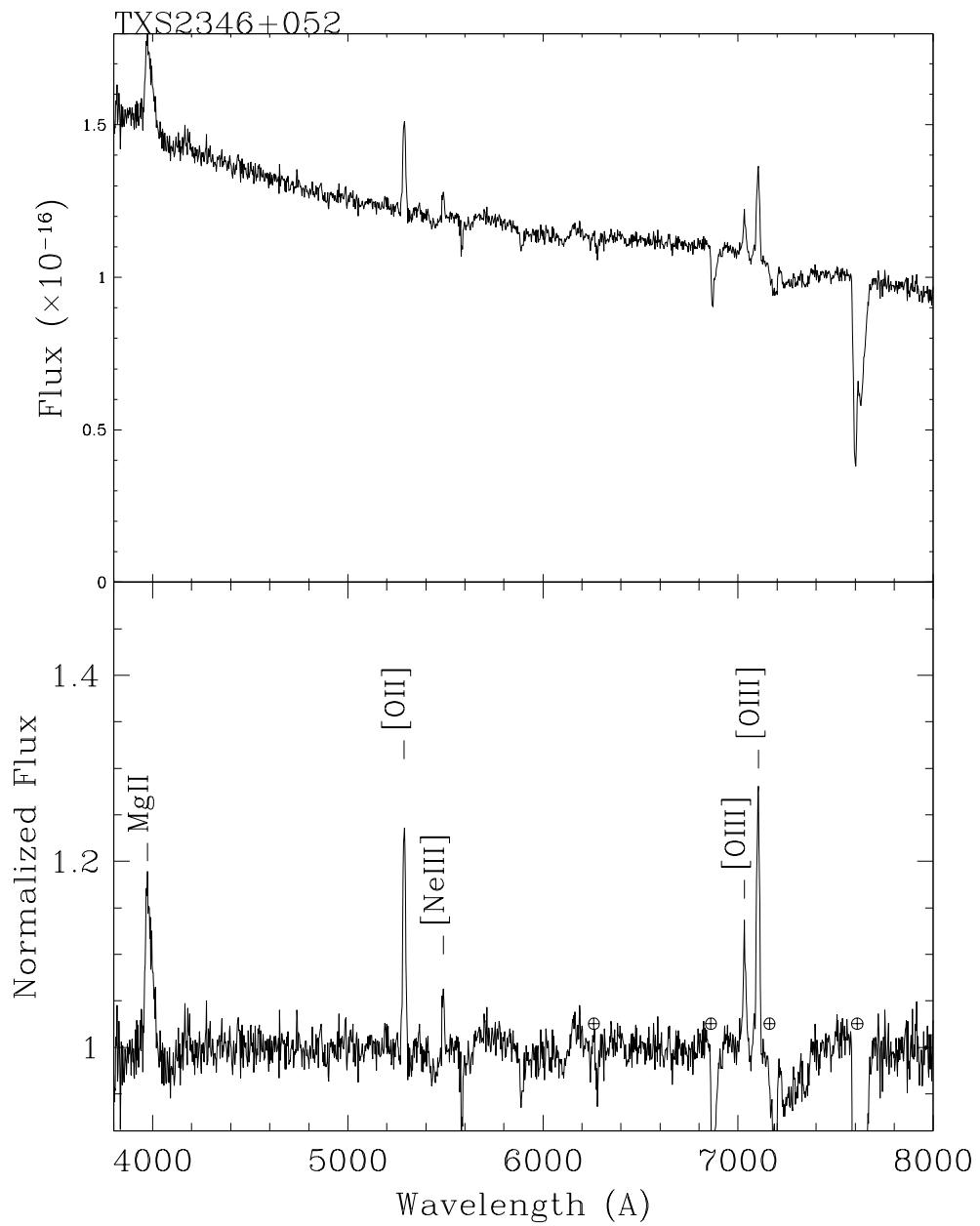


**Figure 1:** –continued.



**Figure 1:** –continued.

POS (BLAZARS 2008) 026



**Figure 1:** –continued.

## References

- [1] B. Sbarufatti, A. Treves, R. Falomo, J. Heidt, J. Kotilainen, R. Scarpa, *ESO Very Large Telescope optical spectroscopy of BL Lacertae. I. New redshifts.* **AJ** **2005** (129) 559
- [2] B. Sbarufatti, A. Treves, R. Falomo, J. Heidt, J. Kotilainen, R. Scarpa, *ESO Very Large Telescope optical spectroscopy of BL Lacertae. II. New redshifts, featureless objects, and classification assessments.* **AJ** **2006** (132) 1
- [3] B. Sbarufatti, S. Ciprini, J. Kotilainen, R. Decarli, A. Treves, A. Veronesi, R. Falomo, *ESO VLT optical spectroscopy of BL Lacertae objects. III. An extension of the sample.* **2008 AJ in press.**

POS (BLAZARS2008) 026