# 16th Microlensing Season of the Optical Gravitational Lensing Experiment

#### A. Udalski Warsaw University Observatory

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OGLE: The Optical Gravitational Lensing Experiment (1992 - ....)

http://ogle.astrouw.edu.pl
http://bulge.princeton.edu/~ogle

#### Three Phases of the OGLE Project

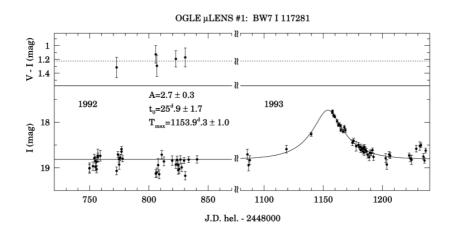
- OGLE-I (1992-1995). 1 m Swope telescope at LCO. ~2 million stars observed. Microlensing
- OGLE-II (1997-2000). 1.3 m Warsaw telescope. ~40 million stars observed. Variable and non-Variable Stars in GB, MC
- OGLE-III (2001-....). 8k x 8k mosaic CCD. ~200 million stars observed (GB, GD, MC). Extrasolar Planets, Microlensing

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#### Las Campanas Observatory, Chile



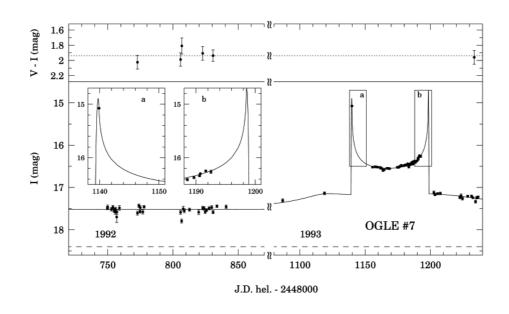
#### OGLE-I Results.



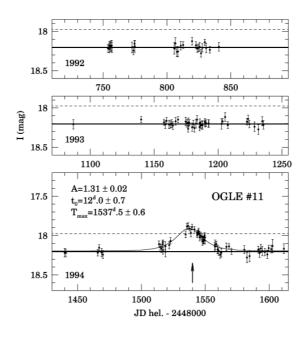
Microlenses: Discovery of the first events toward the GB (1993).

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#### First Binary Microlensing (1994)



#### Early Warning System (EWS – 1994)



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#### Microlensing Optical Depth:

• First empirical determination (1994):  $\tau = 3.3 \times 10^{-6}$ 

#### OGLE-II (1997—2000)

- New telescope new targets: Magellanic Clouds
- New CCD Camera drift scan mode
- ~40 million stars regularly observed
- Variable and non-Variable Stars in GB, MC, gravitational micro and lensing
- Distance scale

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#### 1.3 m Warsaw Telescope at LCO

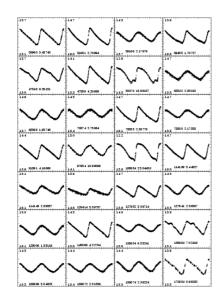


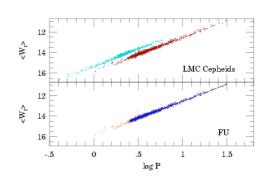
#### Night at LCO



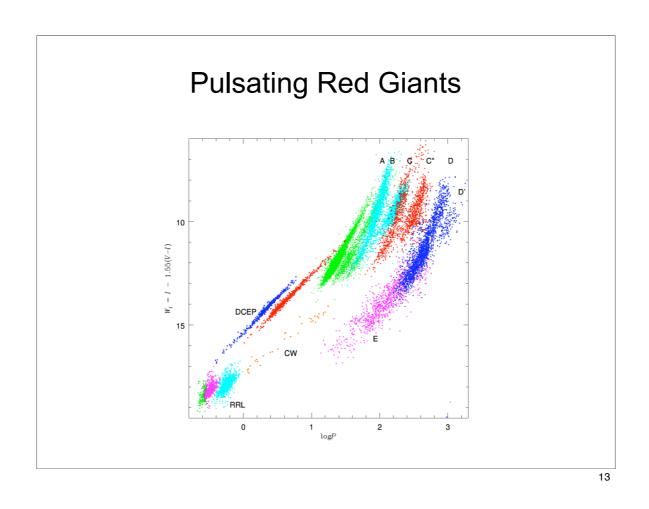
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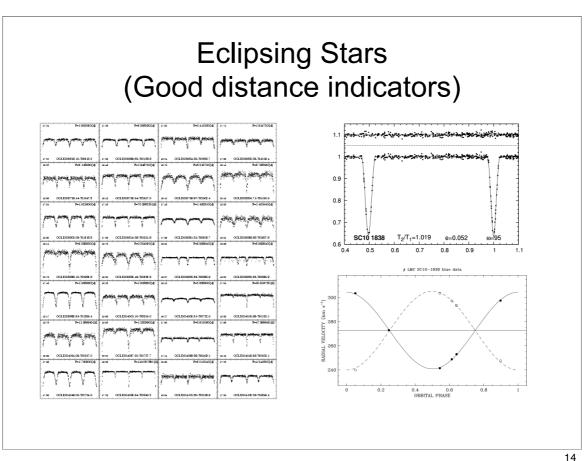
#### Cepheids in the MC. PL relations.



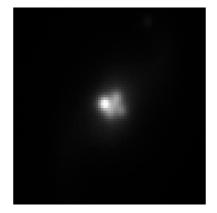


Hubble constant is based on OGLE PL relations





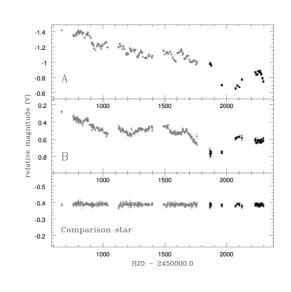
#### Gravitational Lensing. The Einstein Cross.

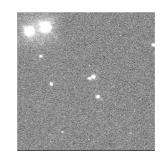


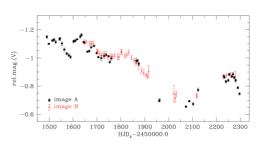


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#### Gravitational Lensing: HE1104—1805







#### OGLE-III (2001-...)

- New 8192 x 8192 pixel mosaic CCD camera (0.26 arcsec/pixel scale): 0.5 x 0.5 sq. degree
- 1.3 m OGLE telescope at Las Campanas Observatory, Chile
- Data Pipeline: photometry derived with image subtraction method (accuracy up to 3 mmag for the brightest stars over a few months long observing run)
- OGLE back in operation on June 12, 2001
- ~200 million stars regularly observed (GB, GD, MC)
- Extrasolar planets, low luminosity objects

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## Warsaw Telescope and 8192x8192 pixel Mosaic Camera



#### Typical OGLE-III Microlensing Season

- EWS Alert System: ~600 microlensing detections (~20 during OGLE-I all four seasons; ~60 per season in OGLE-II)
- EEWS Alert System: ~25 real time anomaly detections
- Several high magnification events, several binary microlensings
- ~<2-3 planetary microlensings</li>

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## Planetary Microlensing – Observing Strategy

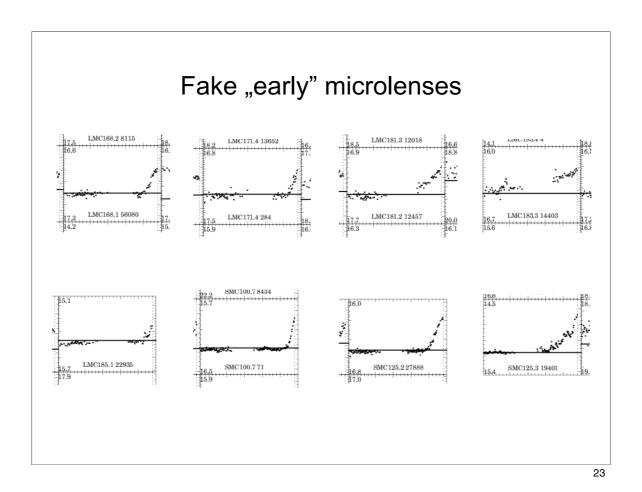
- Detection of microlensing events in early phases by the large survey programs like OGLE or MOA
- Follow-up of the most promising events by followup programs like microFun, PLANET/ROBONET and survey programs in follow-up mode (OGLE, MOA)
- OGLE real time alert system (EEWS) detecting anomalies in the light curve of microlensing events (enabling fast switch to from survey to follow-up observing mode)

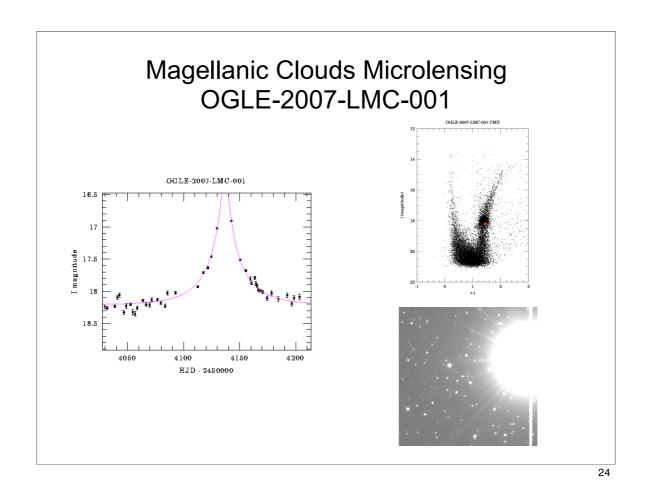
#### 2007 Microlensing Season

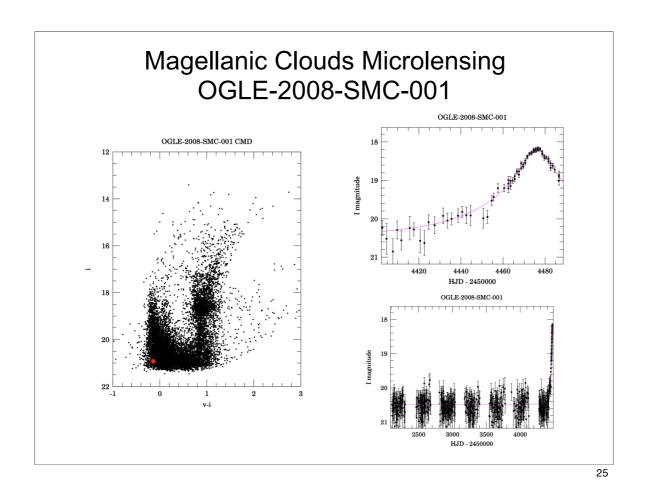
- EWS Alert System: ~600 microlensing detections
- EEWS Alert System: ~20 real time anomaly detections
- Several high magnification events, several binary microlensings

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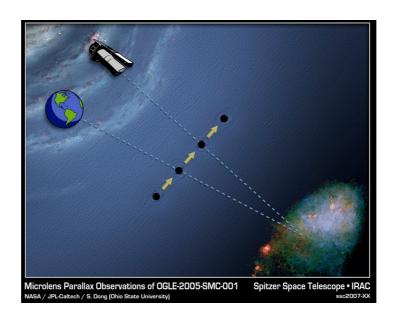
## Magellanic Cloud Microlensing

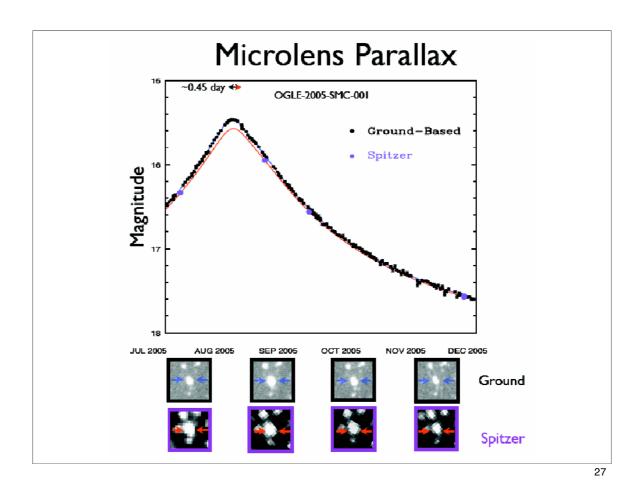


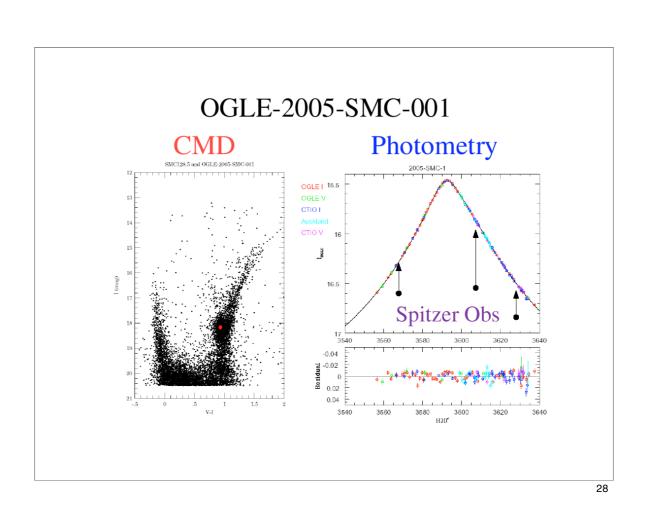




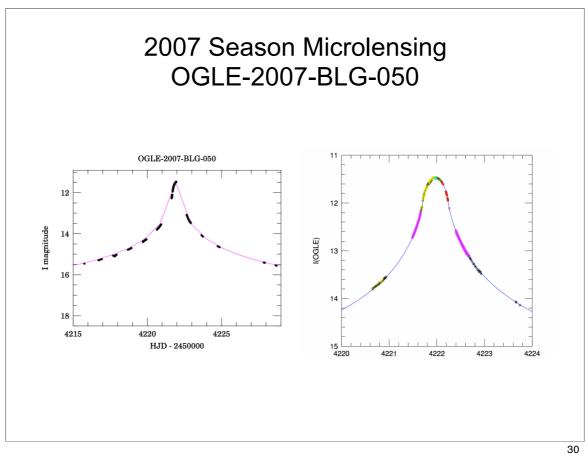
#### OGLE-2005-SMC-001



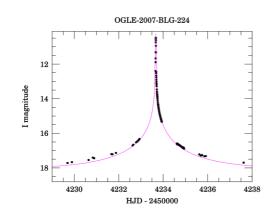


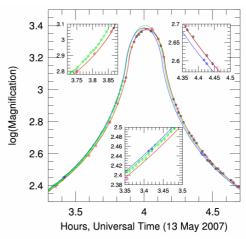


## 2007 Galactic Bulge Microlensing Season



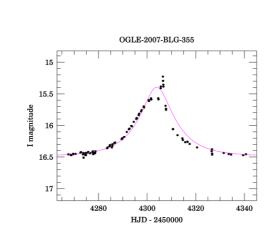


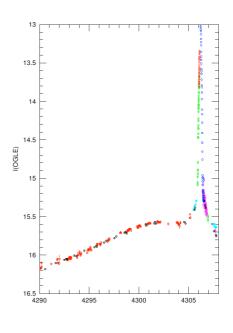




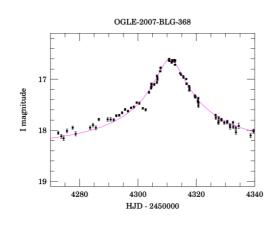
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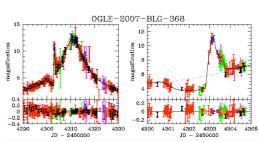
#### 2007 Microlensing Season: Spectacular Event OGLE-2007-BLG-355







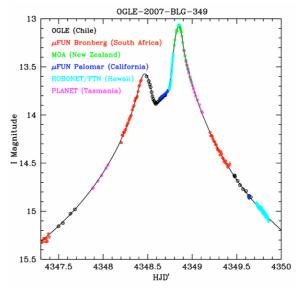




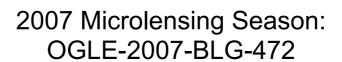
Planetary microlensing.

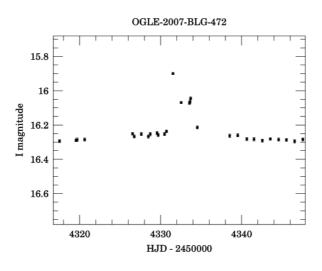
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## 2007 Microlensing Season: OGLE-2007-BLG-349



Planetary microlensing

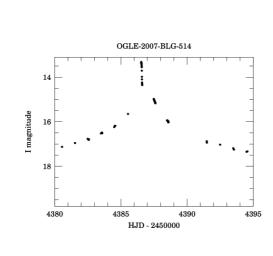




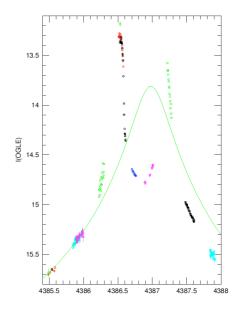
Small q microlensing.

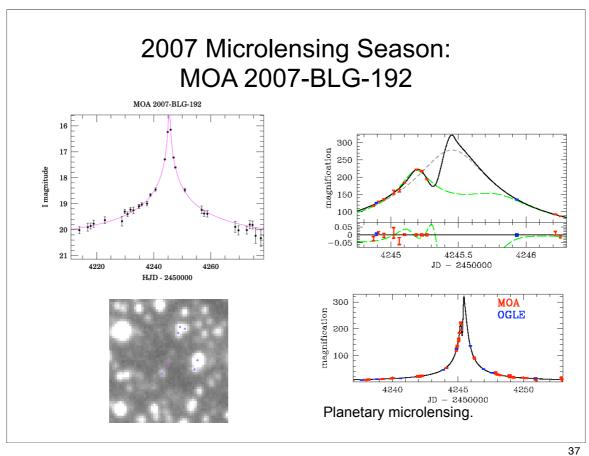
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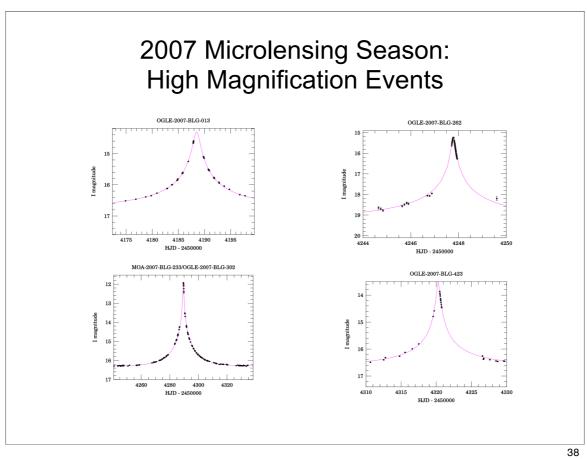
## 2007 Microlensing Season: OGLE-2007-BLG-514

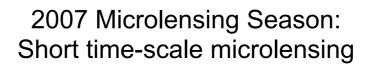


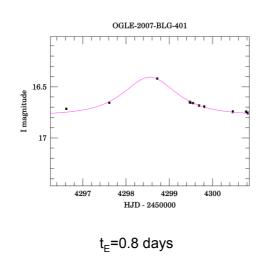
Planetary microlensing?

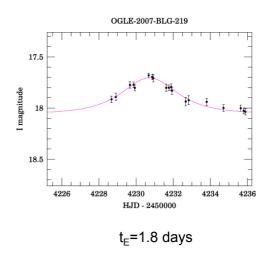






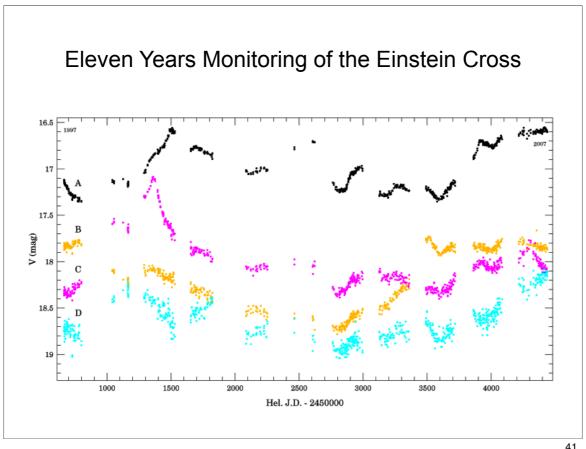






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## Gravitational Lensing Einstein Cross – 2007



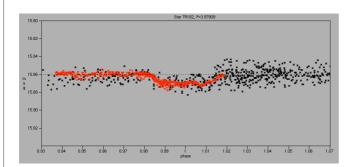
### MISC

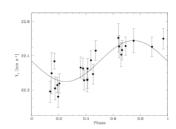
#### OGLE-TR-182

#### A transiting planet among 23 new near-threshold candidates from the OGLE survey\* — OGLE-TR-182

astro-ph/0710.5278

F. Pont<sup>1</sup>, O. Tamuz <sup>2</sup>, A. Udalski<sup>3,4</sup>, T. Mazch <sup>2</sup>, F. Bouchy <sup>2</sup>, C. Melo<sup>3</sup>, D. Nael<sup>9</sup>, C. Moutou<sup>1</sup>/9, R. Diaz<sup>1</sup>, V. Giener<sup>3</sup>, M. Gilon<sup>1</sup>, S. Hoyer<sup>1</sup>, M. Kubiak<sup>3,4</sup>, M. Mayor<sup>4</sup>, D. Mimitig<sup>2</sup>, G. Pietrynssi<sup>3,4,7</sup>, D. Quelou, S. Ramirez<sup>3</sup>, M. T. Ruiz<sup>11</sup>, N.C. Samos<sup>1</sup>, I. Soszyński<sup>3,4</sup>, O. Szewczyk<sup>3,4</sup>, M.K. Szymański<sup>3,4</sup>, S. Udry<sup>1</sup>, K. Harczyki<sup>3,4</sup>, W. Gyrzeniski<sup>3,4</sup>, S. Udry<sup>1</sup>, K.





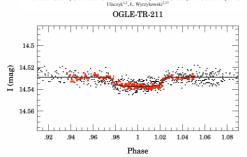
P=3.97910 days, Mp=1.01+/-0.15  $M_J$ , Rp=1.13 +0.24-0.08  $R_J$  Typical hot Jupiter.

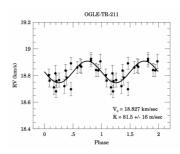
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#### OGLE-TR-211

### OGLE-TR-211 – a new transiting inflated hot Jupiter from the OGLE survey and ESO LP6666 spectroscopic follow-up program \* A. Uddski-1, F. Poni\*, D. Naci\* (, C. Melo\*, F. Bouchy\*, N.C. Samos\*, C. Moutou\*, R.F. Díza\*, W. Gieren\*, M. Gliod\*, S. Hoye\*, M. Mayor\*, T. Mazah\*, D. Mimili\*, G. Fortzyński-2\*, O. Queloz\*, S. Ramine\*, M.T. Rizi\* O. Tamuz-1\*, S. Udry\*, M. Zoccali\*, M. Kuhak\*, M.K. Szymański-2\*, Soxynński-3\*, O. Szewczyk-3\*, K. Ularzyk-1, L. Maryskowski-3\*, M. Szymański-3\*, M. Szymański-3\*, N. Szymański-3\*, O. Szewczyk-3\*, K.

astro-ph/0711.3978





 $P=3.67724 \text{ days}, Mp=1.03 +/- 0.20 M_J, Rp=1.36 +0.18-0.09 R_J$ 

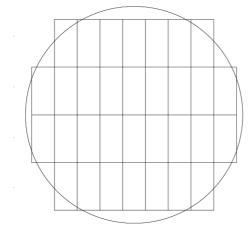
Radius  $\sim$ 20% larger than that of typical hot Jupiter: another "inflated" hot Jupiter.

#### **OGLE 2008 Plans**

- Microlensing campaign 2007 setup
- New mosaic camera with larger field of view

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#### **OGLE-IV**





New OGLE Mosaic Camera – 1.4 square degree field

- Contract with E2V signed 13 chips out of 34 already delivered
- Expected first light Early 2009
- ~10 times larger data flow compared to OGLE-III